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### PART B SOLAR - GEOPHYSICAL DATA

ISSUED FEBRUARY 1960

U. S. DEPARTMENT OF COMMERCE NATIONAL BUREAU OF STANDARDS CENTRAL RADIO PROPAGATION LABORATORY BOULDER, COLORADO



### SOLAR - GEOPHYSICAL DATA

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### SOLAR - GEOPHYSICAL DATA

### INTRODUCTION

This monthly report series is intended to keep research workers abreast of the major particulars of solar activity and the associated ionospheric, radio propagation and other geophysical effects. It is made possible through the cooperation of many observatories, laboratories and agencies as recorded in the detailed description of the tables and graphs which follows. The report is prepared in the Radio Warning Services Section, edited by Miss J.V. Lincoln.

### I DAILY SOLAR INDICES

Relative Sunspot Numbers -- The table includes (1) the daily American relative sunspot numbers,  $R_A$ , as compiled by the Solar Division of the American Association of Variable Star Observers, and (2) the provisional daily Zürich relative sunspot numbers,  $R_Z$ , as communicated by the Swiss Federal Observatory. Because of the time required to collect and reduce the observations,  $R_A$  will normally appear one month later than  $R_Z$ .

The relative sunspot number is an index of the activity of the entire visible disk. It is determined each day without reference to preceding days. Each isolated cluster of sunspots is termed a sunspot group and it may consist of one or a large number of distinct spots whose size can range from 10 or more square degrees of the solar surface down to the limit of resolution (e.g. 1/8 square degrees). The relative sunspot number is defined as R=K(10g+s), where g is the number of sunspot groups and s is the total number of distinct spots. The scale factor K (usually less than unity) depends on the observer and is intended to effect the conversion to the scale originated by Wolf. The observations for sunspot numbers are made by a rather small group of extraordinarily faithful observers, many of them amateurs, each with many years of experience. The counts are made visually with small, suitably protected telescopes.

Final values of  $R_Z$  appear in the IAU <u>Quarterly Bulletin on Solar Activity</u>, the <u>Journal of Geophysical Research</u>, these reports, and elsewhere. They usually differ slightly from the provisional values. The American numbers,  $R_A$ , are not revised.

Solar Flux Values, 2800 Mc -- The table also lists the daily values of solar flux at 2800 Mc recorded in watts/M2/cycle/second bandwidth (x 10-22) in two polarizations by the National Research Council at Ottawa, Canada. These solar radio noise indices are being published in accordance with CCIR Report 25 that a basic solar index for ionospheric propagation should be measured objectively and "preferably refer to a property of the sun such as radiation flux which has direct physical relationship to the ionosphere."

Graph of Sunspot Cycle -- The graph illustrates the recent trend of Cycle 19 of the 11-year sunspot cycle and some predictions of the future level of activity. The customary "12-month" smoothed index, R, is used throughout, the data being final RZ numbers except for the current year. Predictions shown are those made for one year after the latest available datum by the method of A. G. McNish and J. V. Lincoln (Trans. Am. Geophys. Union, 30, 673-685, 1949) modified by the use of regression coefficients and mean cycle values recomputed for Cycles 8 through 18. Cycle 19 began April 1954, when the minimum  $\overline{R}$  of 3.4 was reached.

### II SOLAR CENTERS OF ACTIVITY

<u>Calcium Plage and Sunspot Regions</u> -- The table gives particulars of the centers of activity visible on the solar disk during the preceding month. These are based on estimates made and reported on the day of observation and are therefore of limited reliability.

The table gives the heliographic coordinates of each center (taken as the calcium plage unless two or more significantly and individually active sunspot groups are included in an extended plage) in terms of the Greenwich date of passage of the sun's central meridian (CMP) and the latitude; the serial number of the plage as assigned by McMath-Hulbert Observatory; the serial number of the center in the previous solar rotation, if it is a persisting region; particulars of the plage at CMP: area, central intensity; a summary of the development of the plage during the current transit of the disk, where b = born on disk,  $\ell$  = passed to or from invisible hemisphere, d = died on disk, and  $\ell$  = increasing, - = stable, \ = decreasing; and age in solar rotations; particulars of the associated sunspot group, if any, at CMP: spot count and the summary of development during the current disk transit, similar to the above. The unit of area is a millionth of the area of a solar hemisphere; the central intensity of calcium plages is roughly estimated on a scale of 1 = faint to 5 = very bright.

Calcium plage data are available through the cooperation of the McMath-Hulbert Observatory of the University of Michigan. The sunspot data are compiled from reports from the U. S. Naval Observatory, and from reports from Europe and Japan received through the daily Ursigram messages.

Coronal Line Emission Indices -- In the table are summarized solar coronal emission intensity indices for the green (Fe XIV at  $\lambda$ 5303) and red (Fe X at  $\lambda$ 6374) coronal lines. The indices are based on measurements made at 5° intervals around the periphery of the solar disk by the High Altitude Observatory at Climax, Colorado, and by Harvard University observers at Sacramento Peak (The USAF Upper Air Research Observatory at Sunspot, New Mexico, under contract AF 19(604)-146). The measurements are expressed as the number of millionths of

an Angstrom of the continuum of the center of the solar disk (at the same wavelength as the line) that would contain the same energy as the observed coronal line. The indices have the following meanings:

 $G_6$  = mean of six highest line intensities in quadrant for  $\lambda 5303$ .

 $R_6 = same for \lambda 6374$ .

 $G_1$  = highest value of intensity in quadrant, for  $\lambda 5303$ .

 $R_1 = same for \lambda 6374$ .

The dates given in the table correspond to the approximate time of CMP of the longitude zone represented by the indices. The actual observations were made for the North East and South East quadrants 7 days before; for the South West and North West quadrants 7 days after the CMP date given.

To obtain rough measures of the integrated emission of the entire solar disk in either of the lines, assuming the coronal changes to be small in a half solar rotation, it is satisfactory to perform the following type of summation given in example for 15 October:

where N is the number of indices entering the summation.

Such integrated disk indices as well as integrated whole-sun indices are computed for each day and are published quarterly in the "Solar Activity Summary" issued by the High Altitude Observatory at Boulder, Colorado. In the same reports are given maps of the intensity distribution of coronal emission derived from all available Climax and Sacramento Peak observations, as well as other information on solar activity, such as maps made from daily limb prominence surveys in  ${\rm H}\alpha$  and notes regarding the history of active regions on the solar disk.

Preliminary summaries of solar activity, prepared on a fast schedule, are issued Friday of each week from High Altitude Observatory in conjunction with CRPL and include solar activity through the preceding day. These are useful to groups needing information on the current status of activity on the visible solar disk, but are not recommended for research uses unless such a prompt schedule of reporting is essential. The same information is included in the subsequent quarterly reports, with extensive additions, corrections and evaluations.

### III SOLAR FLARES

Optical Observations -- The table presents the preliminary record of solar flares as reported to the CRPL on a rapid schedule at the sacrifice of detailed accuracy. Definitive and complete data are published later in the Quarterly Bulletin on Solar Activity, I.A.U., in various observatory publications and elsewhere. The present listing serves to identify and roughly describe the phenomena observed.

Reporting directly to the CRPL are the following observatories: McMath-Hulbert, Wendelstein, Sacramento Peak, Mitaka and Swedish Astrophysical Station on Capri. The remainder report through the URS Igram centers or are available through the IGY World Data Center for Solar Activity in Boulder. Observations are in the light of the center of the II-alpha line unless noted otherwise. The reports from Sacramento Peak, New Mexico (communicated to CRPL by the High Altitude Observatory at Boulder) are from observations at the USAF Upper Air Research Observatory at Sunspot, New Mexico, by Harvard University observers, under contract AF 19(604)-4961.

For each flare are listed the reporting observatory, the date, beginning and ending times, time of maximum phase, the heliographic coordinates in degrees, McMath serial number of the region, duration, the flare importance on the IAU scale of 1- to 3+, observing conditions where 1 means poor, 2 fair and 3 good, time of measurement for tabulated width of  ${\rm H}\alpha$  or tabulated area, measured (i.e. projected) maximum area in square degrees, corrected maximum area in square degrees which equals measured area times secant h where h is the heliocentric angle, maximum effective line-width in  ${\rm H}\alpha$  expressed in Angstroms, and maximum intensity of  ${\rm H}\alpha$  expressed in per cent of the continuous spectrum. The following symbols are used in the table:

D = Greater than F = Approximately E = Less than E = Plus

A final column lists provisionally the occurrence of simultaneous ionospheric effects as observed on selected field-strength recordings of distant high-frequency radio transmissions; a more nearly definitive list of these ionospheric effects, including particulars, appears in these reports after the lapse of a month (see below). All times are Universal Time (UT or GCT). Subflares (importance 1-) are listed by date, time of beginning and their heliographic coordinates. A graph presents intervals for which there were no patrols for flare observations from the observatories whose complete data are published in the table.

Ionospheric Effects -- SID, sudden ionospheric disturbances (and GID--gradual ionospheric disturbances) may be detected in a number of ways: short wave fadeouts (SWF), enhancement of low frequency atmospherics (SEA), increases in cosmic absorption (SCNA), and so forth.

A table lists SWF events that have been recognized on field-strength recordings of distant high-frequency radio transmissions. Under a coordinated program, the staffs at the following ionospheric sounding stations contribute reports that are screened and synthesized at CRPL-Boulder: Puerto Rico, Ft. Belvoir, Va., and Anchorage, Alaska (CRPL Stations: PR, BE, AN); Huancayo, Peru (CRPL-Associated Laboratory: NU); and Ft. Monmouth, N.J., White Sands, N. Mex., Adak, Alaska, and Okinawa (U.S. Signal Corps Stations: FM, WS, AD, OK). McMath-Hulbert Observatory (MC) also contributes such reports. In addition, reports are volunteered by RCA Communications Inc., Marconi Wireless, Netherlands Postal and Telecommunications Services, Swedish Telecommunications, and others; these usually specify times of SWF and the radio paths involved. Through the URSIgrams, reports are available from still other stations as given monthly in the footnotes.

In the coordinated program, the abnormal fades of field strength not obviously ascribable to other causes, are described as short wave fadeouts with the following further classification:

S-SWF: sudden drop-out and gradual recovery

Slow S-SWF: drop-out taking 5 to 15 minutes and gradual recovery

G-SWF: gradual disturbance: fade irregular in either

drop-out or recovery or both,

When there is agreement among the various reporting stations on the time (UT) of an event, it is accepted as a widespread phenomenon and listed in the table.

The degree of confidence in identifying the event, a subjective estimate, is reported by the stations and this is summarized in an index of certainty that the event is widespread, ranging from 1 (possible) to 5 (definite). The times given in the table for the event are from the report of a station (underlined in table) that identified it with high confidence. The criteria for the subjective importance rating assigned by each station on a scale of 1- to 3+ include amplitude of the fade, duration and confidence; greater consideration is given to reports on paths near the subsolar point in arriving at the summary importance rating given in the table.

Note: The tables of SID observed at Washington included in CRPL F-reports prior to F-135 were restricted to events classed here as S-SWF.

A second table lists sudden ionospheric disturbances which have been recognized on recorders for detecting cosmic absorption at about 18 Mc (SCNA) or on recorders for detecting enhancements of low frequency atmospherics at about 27 kc (SEA) together with solar radio bursts at 18 Mc as identified on the SCNA records.

Reports are received either directly or through the IGY World Data Center for Solar Activity at the High Altitude Observatory, Boulder, Colo. The following observatories report SCNA: Rensselaer Polytechnic Institute Observatory, Grafton, N.Y. (RE); McMath-Hulbert

Observatory (MC); Sacramento Peak, N.Mex. (SP); High Altitude Observatory, Boulder, Colo. (BO); University of Hawaii, Makapuu Pt., Hawaii (HA); and the Royal Observatory Edinburgh (ED). All of these except the Royal Observatory Edinburgh also report solar noise bursts observed The SEA reports come from the following: Department of Terrestrial Magnetism, Carnegie Institution of Washington, Station at Derwood, Md. (DE); Dunsink Observatory, Ireland (DU); Royal Observatory Edinburgh (ED); three stations operated by the Netherlands PTT at Hollandia, Dutch West Indies (HO), Nederhorst den Berg, Netherland (NE), and Paramaribo, New Guinea (PA); Panska Ves Observatory near Prague, Czech. (PU); High Altitude Observatory, Boulder, Colo. (BO); Sacramento Peak, N. Mex. (SP); McMath-Hulbert Observatory (MC); University of Hawaii (HA): Neustrelitz (NU): Kuhlungsborn (KU); and a group of American Association of Variable Star Observers located at Brooklyn. N.Y. (A1), Pittsburgh, Pa. (A2), Paterson, N.J. (A3), Powell, Ohio (A4), Ramsey, N.J. (A5), Oshkosh, Wis. (A6), China Lake, Calif. (A7), and Manhattan, Kansas (A8).

These reports are coordinated at CRPL-Boulder. When there is agreement among the various reporting stations on the time (UT) of an event, it is accepted as a widespread phenomenon and listed in the table. Some phenomena are listed, if noted at only one location, if there has been a flare or another type of flare-associated effect reported for that time.

In the table under the type of event the importance of the event is given on a scale of 1 minus to 3 plus. Next there is the index of widespread certainty ranging from 1 (possible) to 5 (definite). The time of beginning, maximum and end of the event in UT is given as reported by the station underlined in the group of observing stations. If the event is an SCNA, a percent absorption figure is given. This absorption is calculated by

$$SCNA \% = \frac{I_n - I_f}{I_n} \times 100$$

where I<sub>n</sub> = noise diode current required to give a recorder deflection equal to that which would have occurred in the absence of a flare, i.e. a value extrapolated from cosmic noise level trend before and after a flare. The previous day's record may be considered if necessary.

and  $I_f$  = noise diode current required to give a recorder deflection equal to the level at the time of maximum absorption.

### IV SOLAR RADIO WAVES

### 2800 Mc Observations

The data on solar radio wave events made in Ottawa. Canada by the Radio and Electrical Engineering Division of the National Research Council (A. E. Covington) at 2800 Mc (10-cm emission) are presented, Near local moon (about 1700 UT) the sensitivity of the radiometer is determined and a mean flux for the whole day calculated. These values are given in a tabular form (see table I-1) in units of  $10^{-22}$  watts/  $M^2/c/s$ . Burst phenomena are measured above this level and are given in terms especially suitable for the variations observed on this frequency. The basis for the classifications is described by Covington -J.R. Astro. Soc. Can. 45, 49, 1951 and Dodson. Hedeman and Covington. Ap. J. 119, 541, 1954. A modification in terminology with a view to simplification has been introduced and consists essentially of the omission of the descriptive word "Single" from the "Single-Simple" and "Single-Complex" classes; in designating the "Single", "Single-Simple" and "Rise and Fall" bursts into a single classification designated as "Simple Bursts" with an appropriate type number: in the addition of the letter "f" to indicate that the burst deviates from the basic pattern by the presence of one or more small fluctuations in intensity; and by the addition of the letter "A" to indicate that the event has another smaller duration event superimposed upon it.

### Simple Burst

Any single burst which rises to one maximum and then decreases to the pre-burst level.

- $1 \underline{\text{Simple 1}}$  -- Simple burst, type 1 (formerly "single"). Bursts of intensity less than 7 1/2 flux units and duration less than 7 1/2 minutes.
- 2 Simple 2 -- Simple burst, type 2 (formerly "single-simple"). Bursts of impulsive nature with intensity greater than 7 1/2 flux units.
- $3 \underline{\text{Simple } 3} \underline{\text{Simple } 3}$  Simple burst, type 3 (formerly "rise and fall"). Bursts of moderate intensity with duration greater than 7 1/2 minutes.
- 4 Post-burst increase -- Postburst level is greater than the preburst level. The gradual return to normal flux may require as long as several hours.
  - 5 Absorption following burst (negative post).
- 6 Complex -- (formerly "single-complex"). A single burst which shows two or more comparable maxima before the activity has declined to zero.

- 7 Period of irregular activity or fluctuations -- Series of overlapping bursts of moderate intensity and duration.
- 8 Group -- Series of single isolated bursts occurring in succession with intensity between the events equal to the level before and after the group.
- 9 Precursor -- A small increase of intensity occurring before a larger increase.

### Great Burst

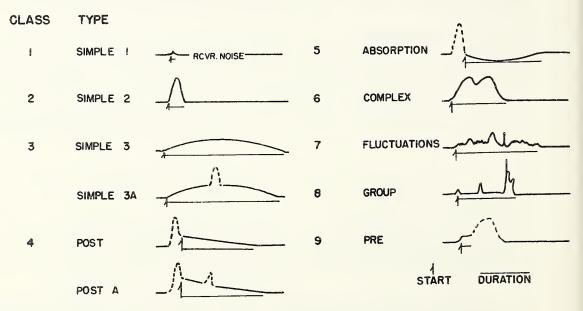
Infrequently occurring bursts of great intensity, often of complicated structure.

### Letter "A"

Indicates that this event has another event superimposed upon it.

### Letter "f"

Indicates that the basic form of the event is modified by secondary fluctuations.



### 200 Mc Observations

Data on solar radio emission on 200 Mc recorded by the University of Hawaii (I. Miyake) at Makapuu Pt., Hawaii, are presented. The outstanding occurrences are reported as described under 170 Mc Observations with the exception that no intensity measurements are given.

### 170 Mc Observations

Data on solar radio emission at the nominal frequency of 170 Mc recorded at the Gunbarrel Hill (Boulder) station of the National Bureau of Standards (C.G. Little) are presented. The half width of the antenna lobe is appreciably greater than the solar disk. Polarization is not determined, but the dipole is oriented E-W. All times are in Universal Time (UT or GCT). Observations are interrupted during the period from 26 to 29 minutes after each hour for calibrations.

Beginning January 1, 1959 the method of reducing the records has been changed. The 3-hourly and daily flux density and variability are no longer determined. The outstanding occurrences are reported. However, instead of giving the intensity to the nearest unit of  $10^{-22}$  watts meter $^{-2}(c/s)^{-1}$ , a scale of 1 to 3 is now used where for the estimate of smoothed maximum flux:

- 1 signifies  $<100 \times 10^{-22} \text{ wm}^{-2}(\text{c/s})^{-1}$
- 2 signifies >100 <1000 x  $10^{-22}$  wm<sup>-2</sup>(c/s)-1
- 3 signifies >1000 x  $10^{-22}$  wm<sup>-2</sup>(c/s)-1.

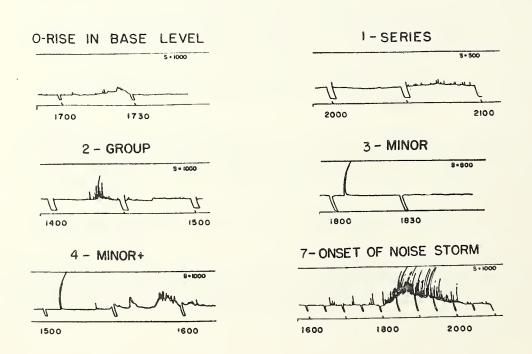
Starting and maximum times are read to the nearest 1/10 minute if they are very definite and otherwise to the nearest minute. If the duration is less than five minutes, it is given to the nearest 1/10 minute; otherwise to the nearest minute. The following qualifying symbols are used:

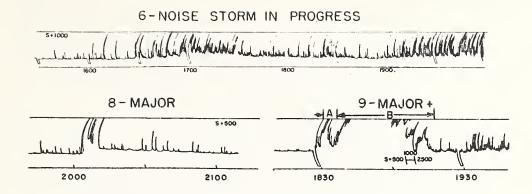
- E = Event in progress before observations began.
- D = Event continues after observations cease.
- I = Event apparently continued during an interruption of the observations. The period of the interruption may be given in the remarks.
- S = Measurement may be influenced by interference or atmospherics.

The types of the outstanding occurrences follow the classification described by Dodson, Hedeman and Owren (Ap J.  $\underline{118}$ , 169, 1953), in which the types are identified by numbers which describe the character of the trace, but not the magnitude of the event, as follows:

- 0 Rise in base level -- A temporary increase in the continuum with duration of the order of tens of minutes to an hour.
- l <u>Series of bursts</u> -- Bursts or groups of bursts, occurring intermittently over an interval of time of the order of minutes or hours. Such series of bursts are assigned as distinctive events only when they occur on a smooth record or show as a distinct change in the activity.

- 2 Groups of bursts -- A cluster of bursts occurring in an interval of time of the order of minutes.
- 3 Minor burst -- A burst of moderate or small amplitude, and duration of the order of one or two minutes.
- 4 Minor burst and second part -- A double rise in flux in which the early rise is a minor burst.
- 6 Noise storm -- A temporary increase in radiation characterized by numerous closely spaced bursts, by an increase in the continuum, or by both. Duration is of the order of hours or days.
- 7 Noise storm begins -- The onset of a noise storm occurs at some time during the observing period.
- 8 <u>Major burst</u> -- An outburst, or other burst of large amplitude and more than average duration. A major burst is usually complex, with a duration of the order of one to ten minutes.
- 9A, 9B, or 9 -- Major burst and second part or large event without distinct first and second parts -- If there is a double rise in flux, the first part, a major burst, is listed as 9A and the second part as 9B. The second part may consist of a rise in base level, a group or series of bursts, a noise storm. A major increase in flux with duration greater than ten minutes but without distinct first and second parts, is listed simply as 9.





Note: In the present table, the type classifications 0 and 1 are not used; they have been included above only for information.

### 169 Mc Interferometric Observations

The 169 Mc interferometric observations are recorded around local noon at Nançay (Cher), France, (N47°23', E8<sup>m</sup>47's) the field station of the Meudon Observatory.

The main lobes are parallel to the meridian plane: the half-power width is 3.8 minutes in the East-West direction and much larger than the solar diameter in the North-South direction. The main lobes are about 2° apart (Ann. Astrophys. 20, 155, 1957). The records give the strip intensity distribution from the center of the disk to 30° to the West and East.

These daily distributions are plotted on the same chart giving diagrams of evolution (C.R. 244, 1460, 1957). Points of intensity 0.5-0.75-1.0-1.5 and 2.0 times  $10^{-22}$  watts/m<sup>2</sup>/c/s are joined day after day in the form of isophotes. Black dots give the position of the center of the radio spots for each day; a line indicates the width of the recorded lobe pattern when it can be measured with certainty. For each radio spot the smoothed intensity around noon is given in  $10^{-22}$  watts/m<sup>2</sup>/c/s.

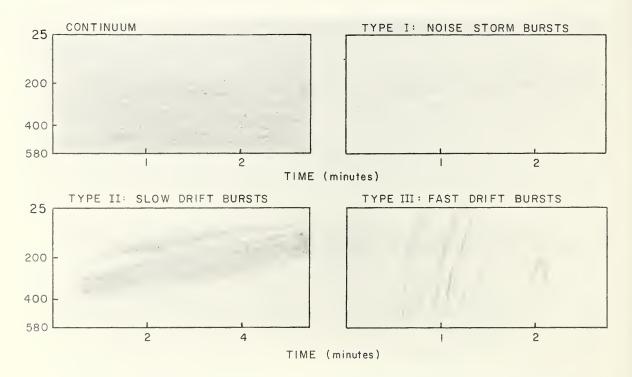
Note that the isophotes cannot be measured when a radio spot of large intensity is on the disk.

### Spectrum Observations

Data on solar radio emission in the spectral range 25-580 Mc/s recorded at the Radio Astronomy Station of Harvard College Observatory, Fort Davis, Texas, are presented. The research program is supported by financial assistance from the Air Force Cambridge Research Center, through the offices of Sacramento Peak Observatory.

The receiving equipment consists of five separate sweep frequency receivers covering the bands 25-50, 50-100, 100-180, 170-320, 300-580 Mc/s. The 25-50 and 50-100 Mc/s receivers are each connected to broad band dipoles which are cross polarised and mounted over a reflecting screen. The other three receivers are attached to separate broad band feeds mounted coaxially at the primary focus of an 8.55 meter diameter paraboloid, the 170-320 Mc/s feed being cross polarised with the other two feeds. The effective collecting area of the antenna is 40 sq. meters at 100 Mc/s and 45 sq. meters at 500 Mc/s.

The four types of recognized spectral activity are idealized below:



The large scale examples of continuum, sometimes called Type IV, are listed as "Cont. IV" in the tables. Photographic examples of the bursts have been published by Maxwell, Swarup, and Thompson (Proc. IRE 46, 142, 1958), and Maxwell (Sky and Telescope 17, 388, 1958; 18, 544, and 556, 1959). A few remaining solar radio bursts are tabulated as unclassified.

The symbols used in the tables are:

b = single burst

g = small group (<10) of bursts G = large group (>10) of bursts

--- = Arrows indicate continuity of solar activity between two Greenwich days.

The minimum detectable level of solar activity is a function of frequency: approximately  $5 \times 10^{-22}$  watts meter  $^{-2}$  (c/s) $^{-1}$  at 500 Mc/s. The equipment records signals over an intensity range of approximately 10,000:1. There are three classes of intensity given in the tables. For 100 Mc/s they are:

1 = Faint, 5 to 40 x  $10^{-22}$  watts meter<sup>-2</sup> (c/s)<sup>-1</sup>.

2 = Moderate, 30 to 200 x  $10^{-22}$ .

 $3 = Strong, >200 \times 10^{-22}$ .

The times are Universal Time (U. T.). The accuracy is to the nearest half minute, except in the case of major outbursts which are specified to the nearest 0.1 minute.

Details of the frequency ranges of activity may be obtained on request to the Radio Astronomy Station, Ft. Davis, Texas.

### **▼** GEOMAGNETIC ACTIVITY INDICES

C, Kp, Ap, and Selected Quiet and Disturbed Days -- The data in the table are: (1) preliminary international character figures, C; (2) geomagnetic planetary three-hour range indices, Kp; (3) daily "equivalent amplitude," Ap; (4) magnetically selected quiet and disturbed days.

This table is made available by the Committee on Characterization of Magnetic Disturbances of IAGA, IUGG. The Meteorological Office, De Bilt, Holland collects the data from magnetic observatories distributed throughout the world, and compiles C and selected days. The Chairman of the Committee computes the planetary and equivalent amplitude indices. The same data are also published quarterly in the <u>Journal of Geophysical Research</u> along with data on sudden commencements (sc) and solar flare effects (sfe).

The C-figure is the arithmetic mean of the subjective classification by all observatories of each day's magnetic activity on a scale of O (quiet) to 2 (storm).

Kp is the mean standardized K-index from 12 observatories between geomagnetic latitudes 47 and 63 degrees. The scale is 0 (very quiet) to 9 (extremely disturbed), expressed in thirds of a unit, e.g. 5- is 4 2/3, 50 is 5 0/3, and 5+ is 5 1/3. This planetary index is designed to measure solar particle-radiation by its magnetic effects, specifically to meet the needs of research workers in the ionospheric field. A complete description of Kp has appeared in Bulletin 12b, "Geomagnetic Indices C and K, 1948" of the Association of Terrestrial Magnetism and Electricity (IATME), International Union of Geodesy and Geophysics.

Ap is a daily index of magnetic activity on a linear scale rather than on the quasi-logarithmic scale of the K-indices. It is the average of the eight values of an intermediate 3-hourly index "ap," defined as one-half the average gamma range of the most disturbed of the three force components, in the three-hour interval at standard stations; in practice, ap is computed from the Kp for the 3-hour interval. The extreme range of the scale of Ap is 0 to 400. The method is described in IATME Bulletin No. 12h (for 1953) p. viii f. Values of Ap (like Kp and Cp) have been published for the Polar Year 1932/33 and for the years 1937 onwards.

The magnetically quiet and disturbed days are selected in accordance with the general outline in <u>Terr. Mag.</u> (predecessor to <u>J. Geophys. Res.</u>) 48, pp 219-227, December 1943. The method in current use calls for ranking the days of a month by their geomagnetic activity as determined from the following three criteria with equal weight: (1) the sum of the eight Kp's; (2) the sum of the squares of the eight Kp's; and (3) the greatest Kp.

<u>Chart of Kp by Solar Rotations</u> -- The graph of Kp by solar rotations is furnished through the courtesy of Dr. J. Bartels, Geophysikalisches Institute, Göttingen.

### VI RADIO PROPAGATION QUALITY INDICES

One can take as the definition of a radio propagation quality index: the measure of the efficiency of a medium-powered radio circuit operated under ideal conditions in all respects, except for the variable effect of the ionosphere on the propagation of the transmittal signal. The indices given here are derived from monitoring and circuit performance reports, and are the nearest practical approximation to the ideal index of propagation quality.

Quality indices are usually expressed on a scale that ranges from one to nine. Indices of four or less are generally taken to represent significant disturbance. (Note that for geomagnetic K-indices, disturbance is represented by higher numbers.) The adjectival equivalents of the integral quality indices are as follows:

1 = useless 4 = poor-to-fair 7 = good 2 = very poor 5 = fair 8 = very good3 = poor 6 = fair-to-good 9 = excellent

CRPL forecasts are expressed on the same scale. The tables summarizing the outcome of forecasts include categories P-Perfect; S-Satisfactory; U-Unsatisfactory; F-Failure. The following conventions apply:

- P forecast quality equal to observed
- U forecast quality two or more grades different from observed when both forecast and observed were > 5, or both < 5
- S forecast quality one grade different from observed
- F other times when forecast quality two or more grades different from observed

Full discussion of the reliability of forecasts requires consideration of many factors besides the over-simplified summary given.

The quality figures represent a consensus of experience with radio propagation conditions. Since they are based entirely on monitoring or traffic reports, the reasons for low quality are not necessarily known and may not be limited to ionospheric storminess. For instance, low quality may result from improper frequency usage for the path and time of day. Although, wherever it is reported, frequency usage is included in the rating of reports, it must often be an assumption that the reports refer to optimum working frequencies. It is more difficult to eliminate from the indices conditions of low quality for reasons such as multipath or interference. These considerations should be taken into account in interpreting research correlations between the Q-figures and solar, auroral, geomagnetic or similar indices.

North Atlantic Radio Path -- The CRPL quality figures, Qa, are compiled by the North Atlantic Radio Warning Service (NARWS), the CRPL forecasting center at Ft. Belvoir, Virginia, from radio traffic data for North Atlantic transmission paths closely approximating New York-to-London. These are reported to CRPL by the Canadian Defense Research Board, Canadian Broadcasting Corporation, and the following agencies of the U.S. Government:--Coast Guard, Navy, Army Signal Corps, U.S. Information Agency. Supplementing these data are CRPL monitoring, direction-finding observations and field-strength measurements of North Atlantic transmissions made at Belvoir.

The original reports are submitted on various scales and for various time intervals. The observations for each 6-hour interval are averaged on the original scale. These 6-hour indices are then adjusted to the 1 to 9 quality-figure scale by a conversion table prepared by comparing the distribution of these indices for at least four months, usually a year, with a master distribution determined from analysis of the reports originally made on the 1 to 9 quality-figure scale. A report whose distribution is the same as the master is thereby converted linearly to the Q-figure scale. The 6-hourly quality figure is the mean of the reports available for that period.

The 6-hourly quality figures are given in this table to the nearest one-third of a unit, e.g. 50 is 5 and 0/3; 5- is 4 and 2/3; 5+ is 5 and 1/3. Other data included are:

- (a) Whole-day radio quality indices, which are weighted averages of the four 6-hourly indices, with half weight given to quality grades 5 and 6. This procedure tends to give whole-day indices suitable for comparison with whole-day advance forecasts which seek to designate the days of significant disturbance or unusually quiet conditions.
- (b) Short-term forecasts, issued every six hours by the North Atlantic Radio Warning Service. These are issued one hour before  $00^h$ ,  $06^h$ ,  $12^h$ ,  $18^h$ , UT and are applicable to the period 1 to 7 hours ahead.
- (c) Advance forecasts (CRPL-J) are issued once a week and are applicable 1 to 7 days ahead. They are modified as necessary by the Special Disturbance Warning (CRPL-SDW) applicable 1 to 3 days ahead, which may be followed by a supplementary forecast (CRPL-Js) applicable to days remaining until next CRPL-J forecast. The forecast entitled "final" consists of the most recent of the above forms and is scored against the whole-day quality index.
- (d) Half-day averages of the geomagnetic K indices measured by the Fredericksburg Magnetic Observatory of the U.S. Coast and Geodetic Survey.

A chart compares the short-term forecasts with Qa-figures. A second chart compares the outcome of the final advance forecasts with a type of "blind" forecast. For the latter, the frequency for each quality grade, as determined from the distribution of quality grades in the four most recent months of the current season, is partitioned among the grades observed in the current month in proportion to the frequencies observed in the current month.

Ranges of useful frequencies on the North Atlantic radio path are shown in a series of diagrams, one for each day. The shaded area indicates the range of frequencies for which transmissions of quality 5 or greater were observed. The blacker the diagram, the quieter the day has been; a narrow strip indicates either high LUHF, low MUF, or both. These diagrams are based on data reported to CRPL by the German Post Office through the Fermeldetechnischen Zentralamtes, Darmstadt, Germany, being observations every one and a half hours of selected transmitters located in the eastern portion of North America. Since January 6, 1958 the transmitters monitored are restricted to those located north of 39° latitude. The magnetic activity index, Apr, from Fredericksburg, Va., is also given for each day.

Note: Beginning with data for September 1955, Qa has been determined from reports that are available within a few hours or at most within a few days, including for the first time, the CRPL observations. Therefore these are the indices by which the forecasters assess every day the conditions in the recent past. Over a period of several years, they have closely paralleled the former Qa indices which excluded CRPL observations and included three additional reports received after a considerable lag. Qa was first published to the nearest one-third of a unit at the same time.

North Pacific Radio Path -- The CRPL quality figures, Qp, are compiled by the North Pacific Radio Warning Service (NPRWS), the CRPL forecasting center at Anchorage, Alaska, from radio traffic data for moderately long transmission paths in the North Pacific equivalent to Seattle-to-Anchorage or Anchorage-to-Tokyo. These include reports to CRPL by the Army Command and Administrative Network, U.S. Air Force and Federal Aeronautical Administration. In addition, there are CRPL monitoring, direction finder observations and field strength measurements of suitable transmissions.

The original reports are on various scales and for various time intervals. The observations for each 8 hours or 24 hour period are averaged on the original scale. This average is compared with reports for the same period in the preceding two months and expressed as a deviation from the 3-month mean. The deviations are put on the 1 to 9 scale of quality which is assumed to have a standard deviation of 1.25 and a mean for the various periods as follows:

07-18 hours UT 5.33 00-24 hours UT 5.67 19-06 6.00

The 12-hour and 24-hour indices Qp are determined separately. Each index is a weighted mean where the CRPL observations have unit weight and the others are weighted by the correlation coefficient with the CRPL observations.

The table, analagous to that for Qa, includes the 12-hourly quality figures; whole day quality figures; short-term forecasts issued by NPRWS twice daily at 06<sup>h</sup> and 18<sup>h</sup> UT, applicable to the stated 12-hour periods; advance forecasts issued weekly by NPRWS (CRPL-Jp report) modified as necessary by Special Disturbance Warnings (CRPL-SDW) and supplementary forecasts (CRPL-Jps); and half-day averages of geomagnetic K indices from Sitka.

The chart compares the outcome of the final advance forecasts, on the same basis as the similar chart for the North Atlantic Radio Path.

Note: Beginning with March 15, 1959 the short-term forecast schedule was changed from three times daily to twice daily. The North Pacific quality figures used for evaluation are now 12-hourly rather than 8-hourly.

### VII ALERT PERIODS AND SPECIAL WORLD INTERVALS

This table gives the Advance Geophysical Alerts as initiated by the Western Hemisphere Regional Warning Center at Ft. Belvoir, Va., and also the Worldwide Geophysical Alerts and Special World Intervals as designated by the World Warning Agency, Ft. Belvoir, Va.

Advance Alerts are of four types, defined as follows:

1 - Solar Flare Alert -- this warning is issued whenever a solar flare of median importance 2 plus or greater has been reported. There will be only one alert issued per flare and only one a day at most.

2 - Magnetic Storm Alert -- this warning is issued whenever a significant magnetic storm, K figure 5 or greater at a middle latitude station

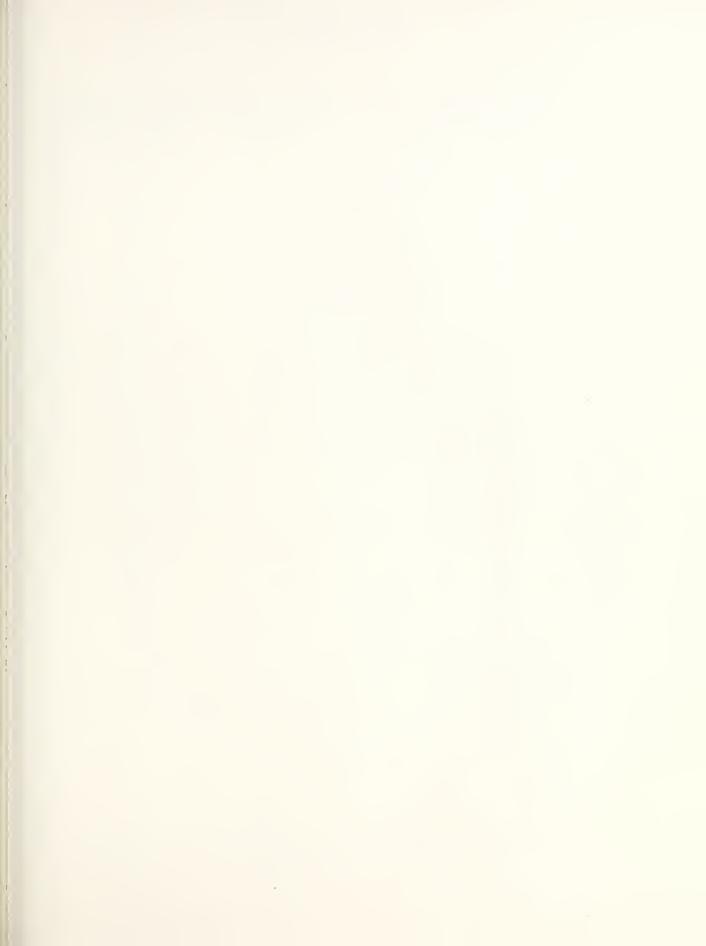
has begun.

3 - Cosmic Ray Alert -- this warning is issued whenever a very outstanding change in cosmic ray flux has been observed -- increase or decrease.

4 - <u>Aurora Alert</u> -- this warning is issued whenever a magnetic storm in middle latitudes has reached K figure 7 intensity or whenever selected auroral stations report the presence of outstanding aurora.

Worldwide Alerts are of the same types as the Advance Alerts, except that the Solar Flare Alert and Cosmic Ray Decrease Alert are omitted. Alert announcements include the event and time of event upon which the alert is based, and, in the case of the Advance Alerts, the station reporting the event.

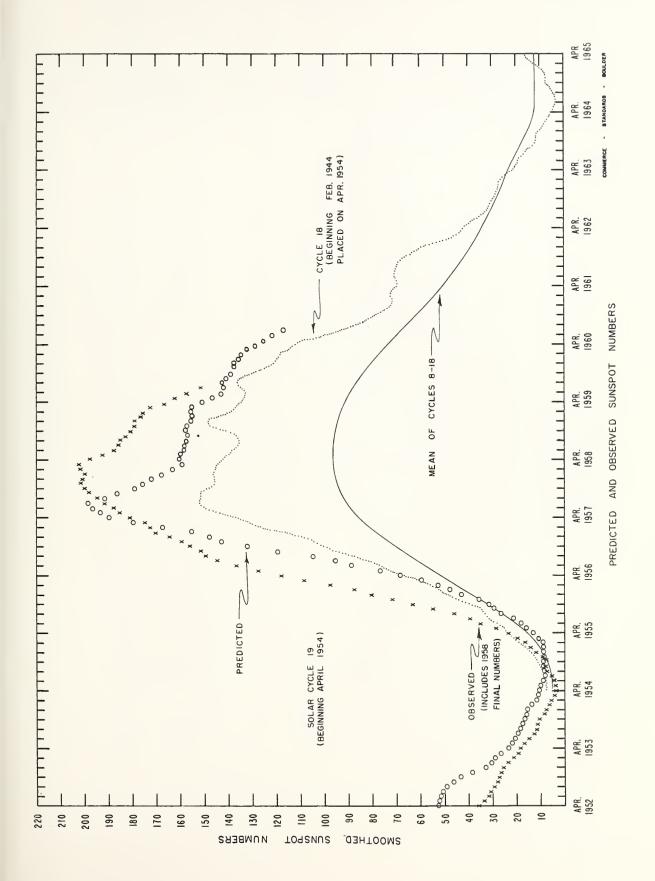
The World Alerts and Special World Intervals are issued by the World Warning Agency on decisions based on Advance Alerts, advice received from Regional Warning Centers and overall policy.



Dec. 1959	American Relative Sunspot Numbers R <sub>A</sub> ,
1	164
2	192
3	194
4	174
5	155
6	152
7	131
8	138
9	97
10	76
11	72
12	74
13	79
14	82
15	120
16	108
17	105
18	131
19	163
20	174
21	154
22	116
23	110
24	121
25	143
26	134
27	150
28	153
29	120
30	113
31	131
Mean:	129.9

Jan. 1960	Zürich Provisional Relative Sunspot Numbers <sup>R</sup> Z	Daily Values Solar Flux at 2800 Mc, Ottawa, Canada Flux
1 2	136 110	175
3 4	133	182
	156	193
5	158	213
6	174	215
<b>7</b> 8	167	224
8	153	219
9	150	201
10	127	194
11	143	200
12	108	184
13	108	178
14	118	176
15	112	183
16	119	183
17	117	179
18	89	176
19	80	164
20	94	157
21	103	162
22	134	172
23	138	188
24	130	210
25	152	<i>2</i> 30
26	209	242
27	186	248
28	159	252
29	193	237
30	178	230
31	178	224
Mean:	139.1	199.7

COMMERCE - STANDARDS - BOULDER



### CALCIUM PLAGE AND SUNSPOT REGIONS

JANUARY 1960

CMP		McMath	Return		Calcium P	lage Data	T	Sunspot	Data
Jan.	Lat	Plage	of		Values			Values	
1960	L	Number	Region	Are	a Int.	History, A	e Are	Count	History
01.1	N09	5512	5478	2000	2.5	$\ell - \ell$ 2	80	7	Ъ/ℓ
02.1	N22	5513	5478	3000	2.5	l\ l 2	50	2	l l l d
03.2	80и	5519	New	400	1	bΛd 1	1 3	_	" "
04.3	S15	5514	5482	3200	2.5	$\ell$ – d 5	80	4	$\ell - \ell$
04.5	N11	5516	New	1400	2.5	l / l 1	20	i	l V l
						_ ~ _	"	•	~ ~ ~
05.6	NO9	5517	5484	700	2.5	$\ell = \ell$ 5	320	6	l - l
06.6	S18	5515	5486	3500	3	$\ell - \ell$ 4	370	3	$\tilde{\ell} - \tilde{\ell}$
07.6	NO7	5521	5495	900	2	$\ell - \ell$ 2	50	2	bΛď
07.7	N25	5520	New	2700	2.5	$\ell - \ell$ 1	20	1	l / l
09.0	S23	5522	New	3200	3	$\ell - \ell$ 1	440	2	$\ell - \ell$
00.1									
09.1	N17	5524	New	600	2.5	$\ell - \ell$ 1	(20)		l √d
09.3	S01	5523	New	2100	3	$\ell - \ell$ 1	560	14	l - l
10.2	S16	5525	New	6300	3	$\ell - \ell$ 1	1020	6	l\ l
10.2	S03	5526	5490	600	2.5	l / l 4	70	3	bΛd
11.6	N20	5527	5491	3500	2.5	$\ell - \ell$ 2	730	5	l - l
12.5	S21	5530	New	500	1	l / l 1	50	1	ь/ г
13.4	NO3	5528	*	3000	2		1 30	T	D / L
14.4	N25	5531	New	500	1.5	$\ell \setminus \ell$ 6,2			
15.1	S16	5529	New	300	1.5	$\ell \setminus d$ 1			
15.9	N10	5532	5497	600	2	$\ell - d  1$			
1 13.9	NIO		3497	800	2	ℓ√d 6			
16.3	S06	5533	5499	800	2.5	l / l 3	50	1	
16.7	N22	5535	New	500	2	$\ell \uparrow \ell$	1 30	1	l \ d
17.7	NO9	5534	5501	2000	2.5	$\ell - \ell$ 4	1		1
17.8	S18	5536	5500	800	1	l / l 4	60	1	
19.3	NO 7	5538	New	2200	3.5	$\begin{pmatrix} \ell & \ell & \ell \\ \ell & -\ell & 1 \end{pmatrix}$	100	5	$\ell \setminus d$ $\ell - \ell$
		3330	1.0%	1 2200	3,3		100	J	k - k
19.5	N20	5537	5502	1700	2	l - l 4			
21.1	N09	5540	**	3 700	3	$\ell - \ell$ 3	150	4	ℓ \ d
21.5	N26	5539	**	2600	3	$\ell - \ell$ 3	540	4	l - L
22.2	N13	5541	**	2900	3	$\ell - \ell$ 3	(20)	(1)	ℓ ¬ d
23.0	S13	5543	New	700	2	$\ell - \ell$ 1	100	3	ρVq
23.3	N26	5542	**	2200	2				
23.4	N11	5545 5545	**	3200 2900	3 3	$\ell - \ell$ 3	200		1
25.0	N11	5546			-	$\ell - \ell$ 3	390	1	l - l
26.5	S13	5547	5507 5510	(5600)	(2.5)	$\ell = \ell$ 10	7.00	,	
26.6	N24	5548	5509	(2000)	(2)	$\begin{pmatrix} \ell - \ell & 2 \\ \ell - \ell & 4 \end{pmatrix}$	190	1	$\ell - \ell$
20.0	1424	40 م	2303	4900	3	$\ell - \ell$ 4	610	5	l _ l
27.2	N09	5549	***	2000	3	$\ell - \ell$ 3	400	10	l - l
29.5	N12	5550	***	11000	3	$\ell - \ell$ 3	580	26	$\tilde{\ell} - \tilde{\ell}$
31.8	S18	5551	5514	5000	3	$\ell - \ell$ 6	510	12	$\tilde{\ell} - \tilde{\ell}$

COMMERCE - STANDARDS - BOULDER

<sup>\* 5493</sup> and 5494. \*\* 5504,5505, 5506. \*\*\* 5511, 5512, 5513.

# CORONAL LINE EMISSION INDICES

### JANUARY 1960

lt (er)	r <sub>1</sub>	24 16 × × ×	x x L x x	77 33 × 69 36 × 64	×××% ×	× × × ×	****	X
North West Quadrant bserved 7 days later)	R <sub>6</sub>	30 13 x x x	* * % * *	41 × 42 × 17 × 17 × 17 × 17 × 17 × 17 × 17 × 1	××× <sup>7</sup> ×	x x 198	* * * * *	X STANDARDS -
ast 7		199 109 x x	167 x x	137 107 107	x 174, x 208	****	****	X COMMENCE - S
North W.	99	161 92 * *	, x 131 , x	113 x 84 82	107 137 x	****	****	×
nt ter)	R1	8,7 8 x x x	× × 2 × ×	36 x	× × × × × × × × × × × × × × × × × × ×	x x 21a x	****	×
South West Quadrant	R,	19 7 x x x	× × 1, × ×	22 × 2 × 2 × 1 2 3 × 2	×××ü×	x x x x x x	****	×
1867	-	124 85 x x	182 x	135 135 89 98	117, 117, 17, 17	****	****	×
South W	95	55 5 × × ×	127 x x	87 87 59 76	52 × 30 ×	****	****	×
int 11er)	R <sub>1</sub>	× × × 57 512	25 x x x x	24 × × × 02	10	16	75 × 50 × × × × × × × × × × × × × × × × ×	×
t (luadrant	R,	××× 42 52	% × × × ×	19 x x x x 2 15	∞ <b>x x x</b> x	x	13 × 13 × 13 ×	×
288 ×	۱,	x x 115 370	190 <b>x</b> x 126	, x , x , x , 78	87 × × × ×	× 7 × × ×	135 x 104 117	261
South Gobserved	95	91 148	146 × × × × 994	× × × 5,2	5 × × × ×	× 7 × × ×	93 61 51 x	157
nt lier)	7.	36 36 24	****	33 40 × × × 40	c ××××	×69×××	36 83 54 x	×
Quadrai	R <sub>6</sub> R	20 25 15	5 ××××	24 × × × × × × × × × × × × × × × × × × ×	17 × × × ×	* 52 * * *	21 40 30 x	×
Eas 7	,,	x x 120 167	194 x x x 102	132 x 114 128	107	156 x x	138 x 152 178	267
North (observed	99	93 141	162 x x 88	711 * * 100 110	8 * * * *	124* ×	123 132# 143	219
CMP Jan	1960	10m4n	6 8 9 10	12521	16 17 18 19 20	55 53 53 53 54 53 53	26 27 28 30	31

a = index computed from low weight data.

x = no observations.

\* = yellow line observed.

SOLAR FLARES
JANUARY 1960

PROVISIONAL	IONOSPHERIC	EFFECT						S-SWF				Slow S-SWF	G-SWF		Slow S-SWF	Slow S-SWF	
	MAX.	. %				30			20 20 30 19	13		30	17	10		20	
	MAX.	WIDTH На		-											6	06.00	
MEASUREMENTS	CORR.	AREA Sq. Deg.		000		6.70	2.50	13 • 00			5.00			3.40	12.00		3.00
	MEAS.	AREA Sq. Deg.	1.40		2.10 1.10	1.90 1.65 1.10 1.90 1.90	2.50	3.70	2.50 2.50 6.40 6.85 3.10	2.76		22.00 10.30 1.90 .50	2.60	1,70	4.80	2.00	
	TIME	TD	0158 1618		1947	0852 0848 2004 2029 2030 2149	0920	1514	1825 1825 1840 1850		· · · · · · · · · · · · · · · · · · ·	2128 2124 2250 2302		1217	1401	1658	
OBS.			3		mm	187181	mm	ммм	NNNN	2	m <b>m</b>	пммм	en.		0,0	7	
ż	POR-	TANCE			16	26 16 1		3 7 7	22211	-	16	3 1 1	1		221	717	-
DURA.	1	MINUTES	32 D 43	. 35 D 11 D 13 D	26 D 4	39 7 18 D 33 22 25	4 D	10 D 38 D	120 120 170 164 U 148 D	24	4 18 D 31 D	195 D 148 D 98 14	54	16 D 63	26 D 79 D	45 45 8	16 D
	McMATH	PLAGE	5507 5520	5511 5520 5515	5513 5525	5525 5525 5512 5512 5512 5512	5517 5525 5525	5525 5512 5512	5515 5515 5520 5520 5520	5527	5515 5516 5516	5527 5527 5527 5527	5525	5525 5530	5525	5533 5533 5540	5539
LOCATION	EOX.	MER. DIST.	¥54 E80	W20 E64 E55	W28 E85	E 77 E 77 W 38 W 37 W 57	E04 E65 E58	E40 W71 W78	W33 W33 W12 W14	£21	W50 W90 W85	E03 E02 ¥08	W37	W 444 W 08	W66 W71	_	E73
90.8	AFF	LAT.	N12 N24	N17 N13 S21	N17 S23	\$19 \$14 NO8 NO9 NO9	N10 S11 S12	S12 N08 N07	\$23 \$23 N28 N26 N26	N 19	\$17 N08 N08	N23 N22 N18 S12	\$10	\$18 \$03	\$20 \$20	806 806 806	N26
		MAX. PHASE	0158		1948	2029 2030 2149			1810 1825 1840 1842	2112	•	2128 E 2124 2250 2302	1650	2305		1658 1732	
OBSERVED	UNIVERSAL TIME	END	0200 D 1650	1025 D 1040 D 1333 D	2010 1952	0906 0854 2014 2058 2048 2207	0959 D 1007	0819 D 1542 D	1950 1950 2050 2050 2048	2130	0815 0925 D 1210 D	2355 D 2326 2422 2312	1710	1232 D 2350	1400 D 1455 D	v O 00	1021 D
	П	START	0128 1607	0950 E 1029 1320	1944 E 1948	0327 0847 1956 E 2025 2026 2142	0920 E 0955 0955	0809 E 0819 E 1504 E	1750 1750 1800 1806 1820 E	2106	0811 0907 E 1139 E	2040 U 2058 E 2244 2258	1646	1216 E 2247		1645 1730 D	1005
DATE	NA	1960	01	02 02 02	03	44444	0.5	07 07 07	8 8 8 8	60	10	1111	12	13	15	15	16
	OBSERVATORY		HAWA I I MCMATH	WENDEL WENDEL WENDEL	HAWAII HAWAII	CAPRI S ARCETRI HAWAII [LOCKHEED [HAWAII	ARCETRI { ARCETRI { WENDEL	ARCETRI ARCETRI CAPRI S	LOCKHEED LOCKHEED LOCKHEED SAC PEAK HAWAII	SAC PEAK	ARCETRI ARCETRI WENDEL	LOCKHEED HAWAII HAWAII HAWAII	SAC PEAK	CAPRI S LOCKHEED	WENDEL CAPRI S	LOCKHEED	WENDEL

SOLAR FLARES
JANUARY 1960

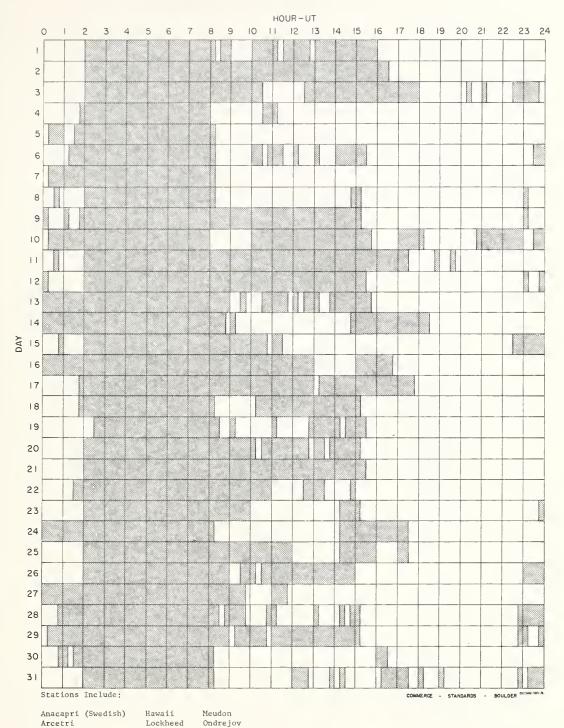
TENCIBLICAGE	CONCENTRAL	EFFECT	Slow S-SWF	Slow S-SWF		G-SWF										
	MAX	INT.	0 7	30		17							16	16	20	30
	MAX	WIDTH На						2.90 2.70 4.80	3.20	3.20	2.30					
MEASUREMENTS	CORR	AREA Sq. Deg.					3 • 00		00 *9	14.00			3.10	5 • 00	2.20	3.00
M	MEAS.	AREA Sq. Deg.	1.40 5.00 1.50	1,0 10 4,0 20 4,0 20 2,0 20	1.00	3.97	1.30	1.90		4 • 00	1.40		2.60	5.00	2.00	2.00
	TIME	T D	2112 2252 2248	0036 1720 1720 1810	2308	1950	2302	0002 1041 1056 1122 2038	0858 0938 1245	1313 1321	1243		1530	0832	1040	0015
OSS	COND.		222	2 1 2 2	ттт	1	6	~~~~N	ммм	mN	23	2	2 2	1 2	1 2	2 2 2
Ä	POR-	TANCE	1 2 1	1 16 16		1	-1	1116	1 1 1 1 2	2 2 16	11		7.7	16		
DURA-	TION	MINUTES	4 D 56 18 D	14 0 112 0 112 0 56 0	9 D 68 88	60 88 D	68 D 16	14 4 D 8 D 13	9 13 D 27 D 18 D	52 D 26 D 50 D	22 28	Q 9	26 24 D	86 D 42	7 D 32 42	28 8 D 27 D 23 D 5 D
NO	McMATH	PLAGE	5539 5540 5540	5545 5545 5545 5542	5541 5545 5541	5545	5528 5541	5539 5538 5538 5538 5549	5550 5538 5550 5550	5549 5549 5549	5550	5550	5549	5550	5551 5550 5554	5550 5550 5551 5551 5551
LOCATION	APPROX.	MER. DIST.	E 62 E 76 E 76	E62 E62 E62 E72	E41 E58 E49	E46	7 W57 2 E25	W W W W W W W W W W W W W W W W W W W	шзшш	E38 E35	E55 E50	E31	E12	E12	E27 W17 E58	3 W22 0 W14 5 E21 7 E17 8 E49
-	Ap	LAT.	N26 N11 N13	N11 N17 N17 N23	N08 N17 N16	N14 N18	NON	N 0 9 N 1 2	N N N N N N N N N N N N N N N N N N N	00 N N 00 00 00 00 00 00 00 00 00 00 00	N06	N05	N01	N07	\$15 N06 \$19	N03 N10 S15 S17 S18
	ш	MAX. PHASE	2112 2252 2248	0036 1720 1750	2308	1944 1950 U	2302	0002 1122 2038	0858		1243		1528	2044	1746 1846	0015
OSSERVED	UNIVERSAL TIME	END	2112 D 2335 2258 D	0046 D 1900 U 1900 U 1856	0943 D 2358 2442	2028 . 2115 U	1146 D 2316	0014 1044 D 1104 1130 2046		1358 D 1335 D 1402 D	1303	1028 D	1544 1544 D	0954 D 2118	1045 D 1810 1902	0033 0118 0922 D 0926 D 0932 D
		START	2108 2239 2240	0032 1708 E 1708 E 1800 E	0934 E 2250 2314	1928 1947 E	1038 E 2300	0000 1040 E 1056 E 1117 2038	0855 0928 E 1241		1241 2004	1022 E	1518 1520 E	0828 E 2036	1038 E 1738 1820	0005 0110 E 0855 E 0903 E
DATE		1960 1960	16 16 16	17 17 17 17	18 18	19	20	23 23 23	24 24 24 24	24 24 24	25	26	27	28	29 29 29	00000
		OBSERVATORY	HAWAII LOCKHEED { HAWAII	HAWAII LOCKHEED LOCKHEED HAWAII	ARCETRI HAWAII HAWAII	{ SAC PEAK LOCKHEED	WENDEL HAWAII.	HAWAII { ONDREJOV ONDREJOV ONDREJOV HAWAII	ONDREJOV ONDREJOV { ONDREJOV WENDEL	{ WENDEL { ONDREJOV { CAPRI S	ONDREJOV HAWAII	ARCETRI	{ SAC PEAK CAPRIS	CAPRI S SAC PEAK	CAPRI S LOCKHEED SAC PEAK	LOCKHEED HAWAII { WENDEL { ARCETRI WENDEL

### SOLAR FLARES JANUARY 1960

_	_			_												_	_		_		Ts.
PROVISIONAL	CIGARAGONOI	TO LOS TO LA COLO TO L	refrect.																		ALL VALUES IN MAX. INT. COLUMN ARE ARBITRARY UNITS (0-40), NOT PERCENT OF CONTINUOUS SPECTRUM. S. THAN 6 PLUS ATER THAN MINUS ROXIMATE
	MAN	IN THE	×																20	20	COMMERCE . STANDARDS MAX. INT. COLUDAN FIS (0-40), NOT PE SPECTRUM. & - PLUS MINUS - NOT REPORTED
	MAX	WIDTH	На																		S IN MAX. INT UNITS (0-40) JOUS SPECTRUM 6 - PLUS - MINUS - NOT I
MEASUREMENTS	CORR	AREA	Sq. Deg.	3.00	4 • 00	3.00	4.00	4.00	8.00	2.20	3.30	5.00	4.00	3.70		2.10	5.00	3.10			COMMENS TO THAMAND TO BE ARBITRARY UNITS (0-40), NOT PERCEN OF CONTINUOUS SPECTRUM.  - LESS THAN  - CREATER THAN  - REPROXIMATE  - APPROXIMATE  COMMENTED  - NOT REPORTED
ME	MEAS	AREA	Sq. Deg.							2.00	3.00			3,00		2.00	4.80	3.00	2.50	3.20	SAC PEAK: E - LES D - GRE U - APP
	TIME	1	υŢ						-	1420	1310			1436		1031	1451	1600	1924	2047	
OBS.	COND.									m	m			m		-	М	1	1	1	HERSTMC
Ä.	POR-	TANCE		1	-	7	7	1	16	7	7	16	1	1		_	-	_	7	-	JINBUR( VATOR)
DURA-	NOIT	ı	MINUTES	22	16 D	12 D	12 D	30 D	163 D	89 D	14 D	15	16 D	30 D		14 D	40 D	26 D	47	28	MOSCOW - GAISH COYAL OBSERVATORY, EDINBURGH GREENWICH ROYAL OBSERVATORY, HERSTMONCEUX SACRAMENTO PEAK SCHAUINSLAND UNITED STATES NAVAL RESEARCH LABORATORY
NOI	McMATH	PLAGE	REGION	5552	5551	5550	5552	5551	5551	5551	5550	5550	5555	5552	!	_		5551	5551	5551	MOSCOW - GAISH ROYAL OBSERVATO GREENWICH ROYAL SACRAMENTO PEAK SCHAUINSLAND UNITED STATES N
LOCATION	APPROX.	-	DIST.	E36	5 E 20	W21	E 19				3 W27	3 W 26	E39	E35					3 E 02	00M	
_	AP	LAT.		60N	\$15	6 ON	N16	\$16	S17	\$17	N03	N03	N21	N12		N 18	816	216	518	518	MOSCOW-G R O EDIN R O HERST SAC PEAK SCHAUINS USNRL
		MAX.	PHASE																1924	2047	i
OBSERVED	UNIVERSAL TIME	END		1057	1101 D		1152 D	1229 D	1442 D	1524 D	1320 D	1321	1350 D	1502 D			1509 D	1605 D	2000	2103	OF GOOD HOPE
	p	START		1035	1045	1138 E	1140 E	1159 E	1159 E	1355 E	1306 E	1306	1334	1432 E		1028	1429	1539	1913	2035	ANACAPRI - GERMAN ANACAPRI - SWEDISH ROYAL OBSERVATORY, CAPE OF KIEV UNIVERSITY KRASHANAL KRASHNYA PAKHRA LOS ANGELES
DATE		ZYZ	1960	30	30	30	30	30	30			30	30	30		31	31	-	31	31	- GERU - SWE: SERVAT TERSIT AL PAKHR,
		OBSERVATORY		WENDEL	WENDEL	WENDEL	WENDEL	WENDEL	\ WENDEL	[CAPRI S	f CAPRI S	1 WENDEL	WENDEL	CAPRI S		CAPRI S	CAPRI S	CAPRI S	LOCKHEED	LOCKHEED	CAPRI G ANACAPRI - GERMAN CARI S ANACARI - SWEDISH GOOD HOPE ROYAL OBSERVATORY, KIEV* KIEV UNIVERSITY KODAIKNAL KODAIKANAL KRASNYA KRASNYA PAKHRA LOCKHEED LOS ANGELES

LOCKHEED OBSERVATIONS: ALL VALUES IN THE MAXI-MUM INTENSITY COLUMN ARE ARBITRARY UNITS ON A SCALE OF 10 TO 40 - NOT PERCENT OF THE CONTINUOUS SPECTRUM.

### INTERVALS OF NO FLARE PATROL OBSERVATIONS JANUARY 1960



Dunsink

McMath

Sacramento Peak

### SUBFLARES

### Noted as follows: Date-Universal Time-Coordinates

### DECEMBER 1959

SAC PEAK	01 1504 N15 W75	* HAWAII 05 1946 E N11 W04	SAC PEAK 13	1654 N15 E20
LOCKHEED LOCKHEED SAC PEAK SAC PEAK LOCKHEED LOCKHEED LOCKHEED	01 1649 N20 E16 01 1808 N25 W90 01 1808 N15 W90 01 1808 N1 E52 01 1808 N10 E54 01 1808 N1 E54 01 2008 N1 E54 01 2018 N1 E51	LOCKHEED 05 2002 NO7 E50  HAWAII 05 2008 N15 W10  LOCKHEED 05 2016 N10 W05  LOCKHEED 05 2028 N09 E01  SAC PEAK 05 2038 N12 W12  LOCKHEED 05 2038 N12 W12  LOCKHEED 05 2038 N10 W12	LOCKHEED 14 SAC PEAK 14 LOCKHEED 14 LOCKHEED 14 LOCKHEED 14 LOCKHEED 14	2017 S14 E1: 2018 S13 E10 2034 S14 E1: 2034 S14 E1: 2220 N12 E28 2304 S15 W7:
SAC PEAK LOCKHEED HAWAII LOCKHEED SAC PEAK SAC PEAK SAC PEAK LOCKHEED HAWAII LOCKHEED LOCKHEED HAWAII LOCKHEED	01 2018 N24 E18 01 2036 N10 E52 01 2038 E N08 W07 01 2041 N12 E02 01 2046 N10 W03 01 2126 N08 W07 01 2126 N08 W07 01 2126 N09 W11 01 2214 N13 W12 01 2228 N24 W85 01 2256 E N14 W01 01 2306 E N14 W01 01 2306 E N14 W01 01 2306 N09 W12	HAWAII 05 2040 N12 M12 LOCKHEED 05 2058 N09 E00 LOCKHEED 05 2058 N09 E00 LOCKHEED 05 2058 N09 E00 LOCKHEED 05 2118 N06 M12 SAC PEAK 05 2112 N07 M11 LOCKHEED 05 2137 N11 M05 SAC PLEK 05 2137 N11 M05 LOCKHEED 05 2154 N11 M05 LOCKHEED 05 2154 N11 M05 LOCKHEED 05 2250 N11 W12 LOCKHEED 05 2220 N11 W12 LOCKHEED 05 2220 N11 W12 LOCKHEED 05 2220 N11 W15 LOCKHEED 05 2230 N11 W16 HAWAII 05 2320 E N13 M14 HAWAII 05 2320 E N13 M14	HAWAII 15  WENDEL 15  **ARCETRI 15  WENDEL 15  **WENDEL 15  **WENDEL 15  **WENDEL 15  WENDEL 15  WENDEL 15  WENDEL 15  WENDEL 15  LOCAMO 15  WENDEL 15  LOCAMEED 15  LOCKHEED 15  LOCKHEED 15  LOCKHEED 15  LOCKHEED 15	0056 E S08 E3: 0823 E S03 E2: 0847 E N11 E5! 1018 E N23 W01 1119 E N16 E0: 1129 E S03 E2: 1201 E N09 E3! 1320 N20 W02 1340 E S15 E56 1652 S06 E3: 1656 N08 E55 1656 N08 E55 1673 N09 W99
HAWA I I WENDEL SAC PEAK SAC PEAK LOCKHEED LOCKHEED LOCKHEED SAC PEAK LOCKHEED SAC PEAK LOCKHEED SAC PEAK LOCKHEED SAC PEAK LOCKHEED	02 0008 E 503 E53 02 1341 E N11 E38 02 1530 N11 E36 02 1554 S10 W18 02 1554 N07 W14 02 1619 N10 E39 02 1622 N 10 E39 02 1626 N09 W17 01 1626 N09 W17 02 1646 N11 E41 02 1646 N11 E41 02 1646 S15 E74,	LOCARNO 06 0955 N11 W12 LOCARNO 06 1250 N09 W20	LOCKHEED 15	1800 N16 W06 1825 N06 W56 1835 N16 W0 1927 S18 E75 1945 N25 W0 1995 S08 E25 1996 S17 E88 2007 S08 E25 2154 N17 W08 2154 N16 W16 2358 S03 Z16
LOCKHEED LOCKHEED LOCKHEED SAC PEAK LOCKHEED SAC PEAK LOCKHEED HAWAII LOCKHEED SAC PEAK LOCKHEED HAWAII SAC PEAK LOCKHEED HAWAII SAC PEAK	02 1648 515 E74. 02 1658 N08 W19 02 1711 515 E74 02 1712 515 E73 02 1748 N10 W06 02 1754 N10 E34 02 1754 N10 E34 02 1756 E N01 E33 02 1756 E N01 E33 02 1756 E N01 E32 02 1855 E N11 E41 02 1815 E N11 E41 02 1815 E N19 W12 02 1826 E N04 E39 02 1824 N11 W14	SAC PEAK 06 1616 NO7 M21 LOCKHEED 0 6 1618 NO7 M22 SAC PEAK 06 1638 NO9 M11 LOCKHEED 06 1635 NO9 M11 LOCKHEED 06 1635 NO9 M11 SAC PEAK 06 1706 NO9 M11 SAC PEAK 07 1710 NO8 W77 LOCKHEED 06 1714 NO8 W70 LOCKHEED 06 1730 NO9 M12 SAC PEAK 06 1740 NO9 M12 SAC PEAK 06 1740 NO9 M14 SAC PEAK 06 1740 NO9 M14 SAC PEAK 06 1820 N11 W26 SAC PEAK 06 1822 N11 W26 SAC PEAK 06 1822 N14 E04 LOCKHEED 06 1822 N15 E03 LOCKHEED 06 1903 N11 M18	* ARCETRI 16  * WENDEL 16  WENDEL 16  MENDEL 16  MENDEL 16  MCMATH 16  LOCKHEED 16  LOCKHEED 16  LOCKHEED 16  LOCKHEED 16  MAMAII 16  LOCKHEED 16  WENDEL 17	0846 E NOS W66 0953 E NOS W66 0953 E NOS W66 1509 S05 E05 1047 S15 W16 1722 S08 W3: 1751 S16 E05 1838 E S18 E66 1840 E S21 E66 0917 E N11 E30
* LOCKHEEO SAC PEAK LOCKHEEO WENOEL ARCETRI WENOEL CAPRIS CAPRIS MCMATH MCMATH LOCKHEEO MCMATH MCMATH MCMATH MCMATH MCMATH MCMATH MCMATH MCMATH MCMATH	02 1850 E N10 M13 02 2108 N09 W21 03 0011 N10 M16 03 0950 E N19 F77 03 0957 E N10 M29 03 1012 E N12 E26 03 1257 E N20 M25 03 1412 N20 M25 03 1412 N20 M25 03 1516 E N10 W20 03 1545 E N31 E46 03 1545 E N31 E46 03 1545 E N31 E46 03 1630 E N31 E28 03 1630 E N31 E28 03 1630 E N31 E38	SAC PEAK 06 1938 NO8 W21 LOCKHEED 06 1940 NO7 W22 LOCKHEED 06 1944 NO7 W22 LOCKHEED 06 1945 S19 W90 M26 LOCKHEED 06 1954 N11 W22 LOCKHEED 06 1954 N11 W22 LOCKHEED 06 1954 N11 W22 LOCKHEED 06 2100 NO9 W23 LOCKHEED 06 2100 NO9 W23 LOCKHEED 06 2100 NO9 W23 LOCKHEED 06 2100 N13 W25 SAC PEAK 06 2140 N13 W25 SAC PEAK 06 2140 N12 W24 LOCKHEED 06 2125 N09 W14 LOCKHEED 06 2152 N09 W14 LOCKHEED 06 2252 N12 W51 LOCKHEED 06 2302 N11 W27 LOCKHEED 07 W276 N09 W14 N	WENGEL   17	0917 E N11 E30 1018 E N11 E30 1120 S16 E55 1547 N15 W36 11548 N16 W36 11724 S03 W01 11739 S06 W47 118018 S07 W01 1824 S07 W01 1828 S04 W0 1828 S04 W0 1828 S08 W47 1829 S08 W47 2158 N25 W36 2144 S05 W47 2158 N25 W36 2209 S17 E52
LOCK HEED	03 1632 NO9 E27 03 1634 N13 M36 03 1746 N13 M36 03 1746 N13 M36 03 1746 N10 E27 03 1992 N0 E27 03 1993 N0 E24 03 1994 N09 E24 03 1994 N09 E24 03 2006 N09 M39 03 2011 N08 M35 03 2016 N09 E22 03 2042 N08 M39 03 2018 E35 03 2045 N08 E35 03 2045 N08 E35 03 2045 N10 E22 03 2045 N10 E22	* CAPRI S 07 1243 E 110 433  WENDEL 07 1314 E 100 433  SAC PEAK 07 1510 110 423  SAC PEAK 07 1510 110 423  HAWAII 07 1820 E 116 437  HAWAII 07 2212 1825 E 50  WENDEL 08 1154 E 129 236  MEMORTH 08 1425 E 113 445  MEMATH 08 1440 E 117 401  MEMATH 08 1441 1828 E 30  MEMATH 08 1621 1813 443  MEMATH 08 1621 1813 443  MEMATH 08 1621 1813 443	CAPR S 18 SAC PEAK 18 SAC PEAK 18 MCMATH 18 LOCKHEED 18 LOCKHEED 18 LOCKHEED 18 SAC PEAK 18 MCMATHED 18 SAC PEAK 18 HOCKHEED 18 HOCKHEED 18 SAC PEAK 18 HOCKHEED 18 HOCKHEED 18 HOCKHEED 18 HOCKHEED 18 HOCKHEED 18 HAWAII 19	1348 NO2 W15 1538 N23 E61 1546 S18 E4C 1548 S18 E4C 1558 E N21 W45 1550 E N21 W45 1816 N30 W20 1936 N10 E13 1940 E N10 E15 2121 S06 W46 0014 S05 W64
LOCKHEEO  HAWAII  SAC PEAK LOCKHEEO SAC PEAK LOCKHEEO SAC PEAK SAC PEAK	03 2250 U NOS M377  04 0146 NOS 220  04 1910 N12 W48  04 1621 N12 W48  04 1621 N10 E16  04 1624 NOS E13  04 1624 NOS E13  04 1624 NOS E13  04 1640 NOS E13  04 1640 NOS E13  04 1640 NOS E15  04 1840 S16 E51  04 1840 NOS E11	## CAPRI S 09 1317 # 110 #48  **SAC PEAK 09 1317 # 110 #48  **SAC PEAK 09 1317 # 110 #48  **CAPRI S 09 1317 # 110 #48  **CAPRI S 10 1222 E N09 M15  **MCMATH 10 1455 N18 E 20  **LOCKHEEO 10 1549 E N08 W74  **LOCKHEEO 10 1559 N20 W27  **LOCKHEEO 10 1610 N09 W74  **LOCKHEEO 10 1610 N09 W74  **LOCKHEEO 10 1938 S06 E 47  **HAWAII 10 2034 E N13 E68  **HAWAII 10 2034 E N13 E68  **HAWAII 10 2034 E N13 W84  **HAWAII 10 2034 E N13 W84	HAWAII 19 HCMATH 20 **CAPRIS 20	0030 NO2 W27 1425 N 23 E46 1548 N 24 E46 1555 E N10 E32 1740 E N24 E46 1745 E N23 E46 1756 E N23 E46 1956 S01 W24 2010 N 11 W03 2010 N 11 W03 2010 N 11 W03 2010 N 10 E3 0010 N 25 E43 0010 N 25 E43 0010 N 25 E31 0010 N 25 E31
LOCKHEED LOCK HEED LOCK HEED LOCK HEED LOCK HEED LOCK HEED LOCK HEED SAC PEAK LOCK HEED	04 1905 N11 603 04 1915 N13 610 04 1917 N13 610 04 1937 N10 600 04 2008 N11 602 04 2008 N11 610 04 2028 N11 602 04 2028 N11 602 04 2127 N08 609 04 2127 N08 609 04 2138 N11 601 04 2152 N09 608 04 2238 N11 601 04 2238 N11 601 04 2352 N09 608 04 2352 N09 608 04 2353 N11 601 04 2353 N11 600	LOCKHEEO 10 2345 NO7 W90  WENDEL 11 1112 E W22 E53 WENDEL 11 1216 E NO9 W76 LOCKHEEO 11 1655 N2 E48 LOCKHEEO 11 1651 N17 E49 LOCKHEEO 11 1722 N11 W85 LOCKHEEO 11 1724 N2 E71 LOCKHEEO 11 1724 N2 E71 LOCKHEEO 11 2032 N16 E40 SAC PEAK 11 2032 N66 W90 LOCKHEEO 11 2039 N05 W85 *LOCKHEEO 11 218 S02 E69 LOCKHEEO 11 2256 N18 W38 LOCKHEEO 11 2258 N11 W38 LOCKHEEO 11 2258 N11 W38 LOCKHEEO 11 2258 N11 W38	SAC PEAK 20 SAC PEAK 20 SAC PEAK 20 SAC PEAK 20 HAWAII 21 LOCKHEE0 21	0142 E N17 W80 0149 E N21 E N2
LOCARNO LOCARNO LOCARNO SAC PEAK SAC PEAK SAC PEAK SAC PEAK LOCKHEEO LOCKHEEO LOCKHEEO LOCKHEEO LOCKHEEO LOCKHEEO LOCKHEEO LOCKHEEO	05 1300 E N10 W10 05 1335 E N09 W01 05 1435 N14 E01 05 1556 N10 W02 1 N09 E02 05 1622 N15 W02 05 1622 N15 W02 05 1650 N13 W05 105 1702 N07 E50 05 1702 N07 E50 05 1712 N07 W09 05 172 N07 W09	ARCETRI 12 0830 E 520 W12  * ARCETRI 12 0831 E 720 E40  WENDEL 12 0850 E N12 E40  WENDEL 12 0850 E N13 E40  WENDEL 12 0950 E N3 F44  WENDEL 12 1995 E N3 F45  LOCKHEEO 12 1713 N42 E35  LOCKHEEO 12 1976 N18 E32  LOCKHEEO 12 1977 N10 E58  LOCKHEEO 12 1924 N23 E35  LOCKHEEO 12 1934 N20 E35  LOCKHEEO 12 2111 N2* E32  LOCKHEEO 12 2112 N2 E35	LOCARNO 22 MCMATH 22 NCSATH 22 NCSATH 22 SAC PEAK 22 SAC PEAK 22 SAC PEAK 22 MCMATH 22 SAC PEAK 22 SAC PEAK 22 MCMATH 22 SAC PEAK 22 HCMATH 22 SAC PEAK 22 HCMATH 22 SAC PEAK 22 SAC PEAK 22 SAC PEAK 22 SAC PEAK 22 WEMPATH 22 WEMPATH 22	1110 NO7 W08 1502 E N10 NO7 1611 N10 W10 1612 N11 W09 1800 N11 N13 W08 1800 N11 N11 808 N12 W07 1848 N15 W13 2100 N25 E21 2126 N10 W10 2132 E N09 W10 2132 E N09 W10 2132 E N11 W13 20930 E N21 E35
LOCKHEED LOCKHEED LOCKHEED HAWAII SAC PEAK LOCKHEED LOCKHEED LOCKHEED SAC PEAK HAWAII SAC PEAK LOCKHEED	05 1749 N13 W06 05 1754 N12 W01 05 1804 N12 W01 05 1804 N12 W10 05 1804 N13 W09 05 1840 N13 W09 05 1849 N12 W02 05 1919 N13 W03 05 1920 N13 W03 05 1922 N14 W09 05 1922 N14 W11 05 1923 N11 W11	LOCKHEEO 12 2210 N23 632 LOCKHEEO 12 2210 N23 632 LOCKHEEO 12 2240 N10 658 LOCKHEEO 12 2242 N10 658 LOCKHEEO 12 2242 N10 618 LOCKHEEO 12 2242 N10 618 LOCKHEEO 12 2377 N18 615 LOCKHEEO 12 2377 N18 615 LOCKHEEO 12 2371 N18 631 HAWALI 13 0052 N14 631 MCMATH 13 1654 6 N17 620	HUANCAYO 23 HUANCAYO 23 LOCARNO 25 CAPRI S 26 LOCHEEO 26 LOCHEEO 26 LOCHEEO 26 LOCHEEO 26 LOCHEEO 26 LOCKEEO 26	0930 E N21 E35 1530 N07 W52 1620 N07 W19 1111 N17 W28 0832 E N24 W35 1625 E N20 W39 1654 N26 W48 1736 N26 W47 1751 N26 W47 1807 N26 W47

### SUBFLARES

### Noted as follows: Date-Universal Time-Coordinates DECEMBER 1959

*Rated as flar	e of	importance	2 1	by other observatories	s (see	CRPL-F 185 P	art B).					conec	AUE - STANDA	106	BOULDER
CAPRI S	2.8	1415	N 20	E 57							LOCKHEED	31	2351	N12	W67
ARCETR1	2.8	1003 E	N08	W26		HAWAII	2.9	2356		E09	IIAWAH	31	2134	506	W14
	-					HAWAII	29	2222		W56	LOCKHEE0	31	2133		E07
HAWAll	27	2254 E		E29		LOCKHEEO	29	2050		W53	LOCKHEED	31	2128		E50
LOCKHEED	27	2233 U		E70		LOCKHEED	29	2045		E13	LOCKHEED	31	1915		E80
HAWAII	27	2230		E 6 7		HAWAII	29	2022		W54	HAWAII	31	1846		W90
ILAWAH	2.7	2217 E		W28		FOCKHEED	29	2020		W53	1 I AWAH	31	1746		W65
LOCKHEED	27	2211		W27		LOCKHEEO	29	1938		W88	LOCKHEEO	31	1742		W59
HAWAII	27	2026		W63		LOCKHEEO	29	1938		W88	SAC PEAK	31	1742		W61
LOCKHEED	27	1946 E		E 73		LOCKHEED	29	1905		E14	LOCKHEEO	31	1713		E01
HAMAII	27	1942		E 75	-	LOCKHEED	29	1755		w90	LOCKHEED	31	1638		W14
LOCKHEEO	27	1832		w90		HAWAII	29	1748 E		W52	LOCKHEED	31	1555		W61
LOCKHEED	27	1829		W17		SAC PEAK	29	1746		W49	LOCKHEED	31	1550		W20
LOCKHEED	27	1739		W57		LOCKHEED LOCKHEED	29	1726		E16	LOCKHEEO	31	1540 E		7 771
LOCKHEED	27	1724		W06		LOCKHEEO	29	1702 1705		. W52 W90	MCMATH	31	1526 E	44.1.0	W60
SAC PEAK	27	1658		E27		LOCKHEED	29	1625		E47	EUCKITEU	,,,	2147 0	.,,,,,	
LOCARNO	27	1336		E55		HAWAII	29	0004 E		E25	LOCKHEED	30	2149 0		. W09
HAWAII	27	0026 E		E 10 E 5 6				000			HAWAII LOCKHEED	30	1934 2105 U		W50 W68
											HAWAII	30	1906		W50
HAWAII	26	2122 E	515	w 73		LOCKHEED	2.8	2240 E	N12	E38	LOCKHEED	30	1830		W68
LOCKHEEO	26	2119	517	w 75		LOCKHEEO	28	2220		E25	LDCKHEEO	30	1830		W68
LOCKHEED	26	2104	N22	W48		LOCKHEEO	2.8	2216		W12	LOCKHEEO	30	1820 E		W50
LOCKHEED	26	2041	N28	E43		HAWAII	28	2050		E28	LOCKHEEO	30	1820 E		W50
LOCKHEED	26	2011	517	W75		HAWAII	2.8	2012	N12	#4I	LOCKHEEO	30	1552		W90
LOCKHEED	26	1914	428	E44		SAC PEAK	2.8	1804	NOB	W36	LOCKHEEO	30	1546		w57
* HAWAII	2.6	1826 E	N 28	W47		SAC PEAK	2.8	1752	N17	W18	CAPRI S	30	∪824 E		E05

### SOLAR FLARES OCTOBER 1959

PROVISIONAL	IONOSPHEBIC	EFFECT						•		S-SWF	S-SWF	S-SWF		S-SWF	S-SWF	S-SWF	S-SWF	
	MAX.	INT.			71		85				74 71	130						100
	MAX.	width На	2 • 70							3.70	6.≥0	2.20 1.70 1.30	2.10		3.00		3.40	
MEASUBEMENTS	CORR.	ABEA Sq. Deg.		2 • 00	2.00 2.10 2.08	1.00	3.30 4.10 8.00 2.10	2.80	3 50		5.80 3.70 4.20 4.90 4.35	8 • 00	3 • 00	3.00		5.00	1.00	
ME	MEAS.	AREA Sq. Deg.			1.80		1.30 3.60 6.06 1.82	1.00	1.10	)	2.00 1.20 2.00 2.35 2.17	3.85	-	•				
	TIME	I E	1122	1008	1155 1200 2303	1010 1319 1513 1600	0630 0716 0716 0718	1241	1348	1418	1447 2013 2334 0002	0500 0505 0510 0504 0505	1043	1216 1340 1340	1329	1439 1510	1529	0415
OBS.	COND.		9	1	2 2	Nwwn	дд	m r	ν α	ю	8 888	2 22	песе	m 12 m	1 M M	w 0	m m m	
ż	POB-	TANCE		н					. – – ∨	1 1 5	2444	2 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7		118	16	1 6	3	п
DURA-	TION	MINUTES	Q 6	8 D	15 D 14 D 19	10 D 4 D 7 D 11 D	13 D 21 5 D 7 D			36	30 20 20 22 8	10 D 10 D 10 D 42	1-66	65 D 147 D	27	18 D 50	36 9 D 16 D	35 D
N.	McMATH	PLAGE	5401	5389	5401 5401 5397	5401 5392 5401 5405	. 5396 5405 5405 5405	5396	5392 5408 5408	5408	5408 5408 5408 5408	5408 5408 5408 5408	5408 5394 5394	5392 5408 5408	5408	400	5408 5405 5408	5392
LOCATION	APPROX.	MER. DIST.	E 59	W52	E28 E28 E09	E08 W51 E00 E40	W67 E28 E30 E28		E65	E62 F60	E62 W71 E61 E62 E60	E55 E55 E55 E55		E 53	E54	E51 E53	E52 E19 E50	06M
	APP	LAT.	517	N08	\$10 \$11 \$16	\$15 \$09 \$08 N04	NO 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	N06 N30	\$11 N32 N30	N30 N32	N N N N N N N N N N N N N N N N N N N	N N N N N N N N N N N N N N N N N N N	N 30 N 28 N 10 N 10	S 18 N 3 2 N 3 2	N30 N29	N31 N32	N29 N08 N31	\$10
		MAX. PHASE			2303		0716 0716 0718	1254	1348 1430 1420		1447 2013 2334 0002	0504	1104	1337	1329		1500	
OBSERVED	UNIVERSAL TIME	END .	1130	1016	1210 1214 D 2320	1020 1323 1520 1606 D	0643 0732 0719 D 0721	1248 131 <b>6</b> 1341	1355 1445 1448	1445	1445 D 1500 2037 2344 0007	0510 D 0510 D 0510 D 0544 0534	1048 1112 1122 D		355	1457 1540	1534 1538 1545	0450 D
	Þ	STABT	1121 E	1008 E	1155 E 1200 E 2301	1010 E 1319 E 1513 E 1555	0630 E 0711 0714 0714 E		1343 1405 1406	1409	1415 E 1440 2005 2322 2359	0500 E 0500 E 0500 E 0502	1033 1041 1103 1103 E	1210 E 1300 E 1305 E	328	1439 E 1450	1458 1529 E 1529 E	0415 E
DATE		0CT 1959	0.1	03	004	0000	90000	900	90	90	90000	0 7 0 0 7 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0000	070	07	07	07	0.8
	Automateur	OBSERVATORY	ONDREJOV	ZURICH	{ ZURICH GOOD HOPE VOROSHILOV	LOCARNO ZURICH ZURICH LOCARNO	GOOD HOPE GOOD HOPE TASHKENT ABASTUMANI	SOOD HOPE	GOOD HOPE LOCARNO GOOD HOPE	ONDREJOV	STOCKHOLM GOOD HOPE HUANCAYO VOROSHILOV	KODAIKNL KODAIKNL KODAIKNL TASHKENT SYDNEY	CONDREJOV ONDREJOV CAPRI G	CAPRI G LOCARNO CAPRI G	CONDREJOV CONDREJOV	ZURICH	ONDREJOV ZURICH ZURICH	KYOTO

SOLAR FLARES OCTOBER 1959

PROVISIONAL	IONOSPHERIC	1021						S-SWF		S-SWF																			
PBOV	ONOI	1						<i>ν</i>		ν 																			
	MAX	*				45		06	70	75							52						59		4 8				
	MAX. WIDTH	Ha			3.00				2.20			2.80			3.40														
MEASUREMENTS	CORR.	Sq. Deg.	2.60	6.00	8 6	2.62	•	9.50	27.00	5.00	21.00	0	2.00	2.10	00 • 9	3.00	6 • 30	2.50	4.20	3.00	00.9	2.80	2.34	ć	2000	3.60	1.00	1.00	
ME.	MEAS. AREA	Sq. Deg.	2.20	5.00		2.30		3.95	16.52	3.00	11.00			1.10		1.50	. 88	1.40	2.10		2.00	1.80	1,38		2.75	2 9 30			
	TIME	T U	0759	1432	1437	1059	1254	0409	0506	0503	0505	1030	1029	1104	1242	0436	1024	1208	1202	1403	1443 2259	0636	0813		0945	23	1243	334	
OBS.	9		m m v	122	13	en e	n ~	7 -	4 10	1	3 1	. 6		m	N 9	- 2	η N	7	I	2	2		2 2	3	2			7 T	9.0
IM:	POR- TANCE	_	~	2-2-	1 E			2.0	1 7 0	7 7	6 -					п,	18		16		7 7		1 1 &		31.			٠,	7 7
DURA-		MINUTES	19 46 D 17	30 40 D	20 10 D 20 D	10		140 D	107 D		26 D 34		21 D	5 D 12	5 D 15 D	21	00	22 D	6 4	) m	O 9	17 D	21 D		28	29		14 D	10 42 D
N	McMATH	REGION	5401 5405 5405	40	5405 5405 5405	5405	4 4	5401	5401	5401	5401	5418	5401	5401	5401	5401	5420	5401	5401	40	5418	40	4 T	40	5408	5416	41	4 0	5405 5401
LOCATION	APPROX.	DIST.	X 0 2 8 0 5 4 0 5 4 0 5 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6	W 0 0 0	W02 W07	W17	W 108	W37	W49	W 56	W54 E15	E70	M44	W44	w46				W57		E 88	W50	W 20 I	¥51	W12	π 3α Ε 42	E39	×20 ×20 ×20	W63 W85
	LAT.		\$14 NO5 NO5	N03	NON NO 2 NO 3	NOI	N08	\$17	\$16	\$19 \$10	S19 N31	508	516	\$17 \$17	\$17 N21	\$17	N07	S16 S17	\$17	517	504 S16	N05	N L S	N05	N 30	\$20 \$20	520	N 0 0	N04 S16
	MAX.	PHASE	1104	43	1437	1059	1254	0409	0200	0501				1104		0436	1024		1202	)  -	2259	,	0813		č	1239			0440
OBSERVED	UNIVERSAL TIME		0807 1137 1114	1455 1508 D	1450 1441 1452 D	1104	1304	0546 D			0523 D	1034		1035 D 1114	1247 1455	0454		1245 D	1257		1444 D	0653	0815 0904 D	1030	0950	1303	1249	1349	0645
	START		0748 1051 E 1057	1425 1428 E	1430 1431 E 1432 E	1054 1058 F		0326 E	0437 E	0457	0457		1029 E		1242 E 1440 E	0433	1022	1123 E		1400 E	144 <i>2</i> 2258	0636 E	0843 E		0922	1234		1335 E	0635 0713 E
DATE	0CT	1959	0 0 8	0.8	8 8 8	60	60	10	0 0	00	00	00	10	10	10	11	11	11	11	11	11	12	12	12	12	12	12	12	13
	OBSERVATORY		CAPRI S CAPRI G LOCARNO	LOCARNO CAPRI S	MEUDON ONDREJOV CAPRI G	PIRCULI CAPRI G		ALMA-ATA SYDNEY	TASHKENT	SYDNEY	LOCARNO	SUBTE ON STIELL	SURICH	LOCARNO GOOD HOPE	ONDREJOV CAPRI G	SYDNEY		CAPRI S		CAPRI G	SYDNEY	GOOD HOPE	PIRCULI [ PIRCULI	LOCARNO	PIRCULI	GOOD HOPE	ZURICH	ZURICH	LOCARNO { CAPRI G

## SOLAR FLARES OCTOBER 1959

PROVISIONAL	IONOSPHEBIC	EPPECT						S-SWF S-SWF
	MAX.	. K	185	44 44		56	73 80 78 110 61	74
	MAX.	WIDTH Ha	1.40			2,50	3.80 2.90	2.20
MEASUREMENTS	CORR.	AREA Sq. Deg.	11.000 3.30 1.00	1.85 4.00 4.00 1.88	3.00	200 00 00 00 00 00 00 00 00 00 00 00 00		1 000 4 4 4 4 4 4 4 4 4 4 4 4 4 4 1 1 0 0 0 0
M	MEAS.	AREA Sq. Deg.	1.90 1.20 1.10 1.10	1.29 .73 .19 .92	2.60	1,50 1,47 1,20 1,20 1,10 2,72	3.34 3.07 3.07 1.01 1.82 1.82	3.11
9	TIME	۱۵	0743 0821 0852 0838 0858	0533 0600 1246 1149	0722 1036 1100	00557 07057 07057 0855 0855 0855 0858 0858	2218 2218 2343 0121 0608 0611 0819 1117 2256	0627 0707 0710 0808 1025 1322 1322
OBS.	COND.		887888		8 H 8 8	NNW H H NWW	าคค ผพคพค	
Ė	POR-	TANCE	1 1 1 1 6		1 1 1 5	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	
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NO	McMATH	PLAGE REGION	5401 5401 5401 5401 5401 5401 5401	5423 5424 5405 5405 5408 5423 5423	5418 5418 5418 5418 5408	54084 54084 54084 544277 544277 544277 54627	5427 5427 5427 5427 5427 5427	5427 5427 5427 5420 5418 5425 5425
LOCATION	OX.	MER. DIST.	X X X X X X X X X X X X X X X X X X X	W39 W85 W52 W57	W04 W11 W09 W08 W63	XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX	ПП К Б С С С С С С С С С С С С С С С С С С	E32 E31 W22 W50 E11 E11
	APPROX.	LAT.	\$15 \$18 \$16 \$16 \$18 \$18 \$15 \$15 \$19	\$09 \$08 \$08 \$06 \$06 \$06 \$06 \$06 \$06 \$06 \$08 \$08 \$08 \$08 \$08 \$08 \$08 \$08 \$08 \$08	\$05 \$05 \$06 \$05 N29	0	N N N N N N N N N N N N N N N N N N N	N
		MAX. PHASE	082 <b>1</b> 0852	0533 0600 1149		0057 0705 0900 0855 0858 0858	2343 0121 0607 0819 U	1025
OBSERVED	UNIVERSAL TIME	END	0805 0925 0835 0857 D 0800 0900 0910	0549 D 0602 1252 D 1200 D 1245 D	0726 D 1101 D 1120 1100 D 1232	0104 0735 D 0801 D 0920 0930 D 0915 0900 D 0918 D	2320 0001 0132 0617 0615 0830 D 1129	0632 D 0715 0720 0842 D 1027 1356 1420
		START	0.743 E 0.821 E 0.835 E 0.839 E 0.845	0518 E 0558 1245 E 1139 E 1230 E	0721 E 1023 1025 1040 1227 E		2337 2337 0112 0603 0610 E 0815 E 1117 E	0624 0706 0708 0807 1023 1316 1320
DATE		OCT 1959	113333333	14 14 14 15 15	16 16 16 16	7	17 17 18 18 18 18	19 19 19 19 19
	ORSERVATORY		COOD HOPE  GOOD HOPE  CAPRI SIMEIZ SINNEIZ LOCARNO LOCARNO ZURICH	PIRCULI PIRCULI CAPRI G PIRCULI PIRCULI LOCARNO	CAPRI G CAPRI S LOCARNO AROSA ZURICH	SYDNEY PIRCULI ONDREJOV (LOCARNO SIMEIZ GOOD HOPE CAPRI G AROSA PIRCULI CABASTUMANI CAPRI G	VOROSHILOV VOROSHILOV VOROSHILOV TASHKENI CONDREJOV SIMBIZ CAPRI G VOROSHILOV	ONDREJOV  { SIMEIZ
L-								

# SOLAR FLARES OCTOBER 1959

PROVISIONAL	IONOBPHERIC	EFFECT							
Bd	1 .	INT.	120	6 6 3	63 70 76 76 52 52 100 46	140	5 4 5 5 8 8 6 5 5 8 8 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9	5 7 2 2	
	MAX.	WIDTH	1.90 2.30		2.90	3.90		****	
MEASUREMENTS	CORR.	AREA Sq. Deg.	8 • 00 2 • 80	1	2 2 2 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	4 4 7 7 7 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	3.00 1.28 5.00 7.00 7.00	2. 30 1. 12 2. 53	11.00
MEA	MEAS.	AREA Sq. Deg.	2.50	1.000 .900 .46	1.08 1.82 1.09 1.00 2.02 1.82 1.82	2.08	3.00 . 83 . 55	1.80 1.81 1.60 1.01	00•9
	TIME	- n	1351 1400 1405 1407 2335	0630 0853 0844 0846 1215 1239 1333 1417	0800 0800 0800 0814 0835 0956 11003	0658 0710 0801 1021 1241 1249	0313 0734 0744 0750 1123 1525	0802 0752 0755 0758 0804 0828	0150
OBS.	COND.		~ ~ ~ ~ ~ ~	<b>NNMMM</b> MM	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	<u> </u>	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	~~~ ~~~	n
Ė	POR-	TANCE	1 1 2 1 1		1 1 1 6		2 S & & S &		2
DURA-	TION	97	14 18 D 33 D 22 D 23 D 9 D	26 336 60 60 60 60 60 60 60 60 60 60 60 60 60	14 8 D 11 D 29 D 14 6 D 13 D	14 D 23 D 15 D 23 D 23 D	25 D 11 17 20 12 D 18 D 15 D	27 D 6 D 11 21 11 15 20 D	Q 99
	McMATH	PLAGE REGION	5427 5430 5430 5428 5428 5418	00000000000000000000000000000000000000	5418 5418 5418 5427 5427 5427 5418 5418	54438 5433 5438 5438 5431	5421 5421 5431 5433 5433 5433 5433	5431 5431 5431 5427 5427 5433	5428
LOCATION	ox.	MER. DIST.	E26 E27 E21 E22 E90	E85 E16 E81 E78 E75 E08 E15	E E E E E E E E E E E E E E E E E E E	E90 W17 E89 E77 W66 W17	W W Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z	W38 W39 W39 W39 E21	W 50
	APPROX	LAT.	\$06 N08 N09 N17 N08	X X X X X X X X X X X X X X X X X X X	010 000 000 000 000 000 000 000 000 000	S11 S11 S10 S00 N13 N13	S113 N04 N05 N05 N05 S008	\$05 \$08 \$12 \$03 \$02 \$02 \$12	N40
		MAX. PHASE		0853 U	0758 0800 0803 0814 0848 0956 1003	0801	0313 0734 0744 0750	0755 0758 0804 0828	0150
OBSERVED	UNIVERSAL TIME	END	1354 1414 D 1430 1420 1530 D 2344 D	0656 0910 0849 0849 0849 1245 01245 1412 1402 0432	0810 0805 0810 0821 0845 0905 1006 1145 D	0712 D 0715 0821 D 1035 D 1259 1257 D	0327 D 0740 0755 0800 1132 D 1520 1530 D	0814 D 0756 D 0801 0815 0810 0837	0234 D
	D	START	1340 1356 1357 E 1358 E 1507 E 2335 E	0630 E 0837 E 0843 E 1210 E 1237 E 1330 E 1416 E	0756 0758 0802 E 0806 E 0834 E 0836 0952 1000	0658 E 0709 E 0758 E 1020 E 1236 1248 E	0302 0729 0738 0740 1120 E 1502 1517	0747 0750 E 0750 0754 0759 0822	0128
DATE		0CT 1959	19 19 19 19	00000000	21 21 21 21 21 21 21	222222	2222222 2522222	4 4 4 4 4 4 5 5 5 5 5 5 5 5 5 5 5 5 5 5	2.5
	and Carried Control	ОВЗЕНУАТОВУ	t ondrejov C ondrejov C oppri G C oppri S C oppri C	GOOD HOPE SIMEIZ CAPRI G CAPRI G CAPRI G CAPRI G CAPRI G CAPRI G	SIMEIZ KRASNYA ABASTUMANI PONDREJOV PIRCULI PIRCULI KRASNYA KRASNYA CAPRI G CAPRI G	KYOTO CAPRI G TASHKENT CAPRI G CAOR HOPE CAPRI G	SYDNEY PIRCULI PIRCULI PIRCULI CARRI G CARRI G CARRI G CAPRI G	CAPRI S CAPRI G PIRCULI GOOD HOPE PIRCULI PIRCULI CAPRI G	SYDNEY

# SOLAR FLARES

OCTOBER 1959

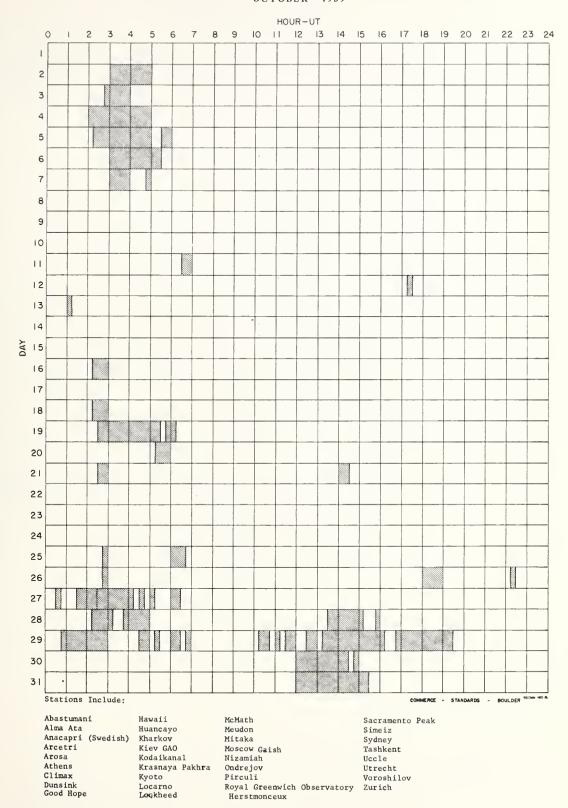
								_		_				_			
PROVISIONAL	EFFECT																
MAX.	INT.	48											100		100	100	115
MAX.	WIDTH							1.90									
MEASUREMENTS CORR.	AREA Sq. Deg.	1,31		4 • 00		1.00	4 • 00		2.80	4 • 00		2.50	1.40	2.00			. 80
MEAS.	AREA Sq. Deg.	1,19							1.20		06.	2.40	2.27	2.00	1.25	3.74	1.40
TIME	T D	0704		1202		1452	0748	1046	0742	0948	1039	1051	1049	0158	0159	0158	0914
OBS.		1	3	9	9	М	3	n		9			1	2			2
IM- POR-	TANCE	1	7	7	7	-	П	-	1	7	7	П	-	-1	П	1	7
DURA. TION	MINUTES	10	30	12 D	7	13	13 D	16 D	11	5 D	18	12	21	21	7 D	0 6	J.
McMATH	PLAGE	5433	5438	5433	5438	5437	5433	5433	5431	5433	5431	5433	5440	5438	5438	5438	5437
LOCATION IOX.	MER. DIST.	E 11	E44	E08	E42	E22	E02	E01	W63	W13	M77	W15	E30	W34	W28	W33	W31
APPROX.	LAT.	N13	808	40N	512	60N	N05	N05	516	N05	\$15	6 ON	N29	510	\$10	808	N11
t:	MAX. PHASE	0704				1500			0742		1039	1051	1049	0158			0914
OBSERVED UNIVERSAL TIME	END	0710	1055	1212	1440	1505	0758 D	1101	0420	0.952 D	1052	1102	1106	0216		0207 D	0919
	START	0020	1025	1200 E	1433	1452	0745	1045 E	0739	0947 E	1034	1050	1045	0155	0158 E	0158 E	0914
DATE	0CT 1959	25	2.5	25	25	25	56	26	2.7	27	27	27	28	31	31		31
	OBSERVATORY	PIRCULI	LOCARNO	CAPRI G	LOCARNO	LOCARNO	CAPRI G	ONDREJOV	GOOD HOPE	CAPRI G	GOOD HOPE	GOOD HOPE	KRASNYA	SYDNEY	KYOTO	(KYOTO	KRASNYA

SAC PEAK: ALL VALUES IN MAX, INT, COLUMN ARE ARBITRARY UNITS (0-40), NOT PERCENT OF CONTINIORS SPECTERM	E - LESS THAN & - PLUS D - GREATER THAN MINUS U - APPROXIMATE \( \Boxed{1}\) - NOT REPORTED
MOSCOW-G MOSCOW - GAISH R O EDIN ROYAL OBSERVATORY, EDINBURGH R O HERST GREEWLICH ROYAL, OBSTRVATORY HTRSTMONCEHX	SACRAMENTO PEAK SCHAUINSLAND UNITED STATES NAVAL RESEARCH LABORATORY
MOSCOM-G M R O EDIN R R O HERST G	SAC PEAK S SCHAUINS S USNRL U
MAN EDISH FORY, CAPE OF GOOD HOPE	LA LA
AMACAPRI - GERMAN ANACAPRI - SWEDISH ROYAL OBSERVATORY, C	KIEV UNIVERSITY KODAIKANAL KRASNAYA PAKHRA LOS ANGELES

CAPRI G CAPRI S GOOD HOPE KIEV\* KODAIKNAL KRASNYA LOCKHEED

COMMERCE - STANDARDS - BOULDER

LOCKHEED OBSERVATIONS: ALL VALUES IN THE MAXI-MUM INTENSITY COLUMN ARE ARBITRARY UNITS ON A SCALE OF 1 TO 4 - NOT PERCENT OF THE CONTINUOUS SPECTRUM.



#### IONOSPHERIC EFFECTS OF SOLAR FLARES

(SHORT-WAVE RADIO FADEOUTS)

#### DECEMBER 1959

Dec.	Start	End	Type	Wide	Impor-	Observation Stations	Known
1959	UT	UT	· ·	Spread	tance		Flare, UT
				Index			CRPL-F 185
1	1247	1327	S-SWF	1	3	JU	1208E
1	1512	1613	Slow S-SWF	5	2+	BE, FM, HU, LA, MC, NE, PR, WS, CW* *	1456E
1	1705	1900	S-SWF	5	3	AN, BE, FM, HU, LA, MC, NE, PR, WS, CW***	1638
2	0330	0450	S-SWF	1	3-	<u>OK</u>	*
2	0500	0550	Slow S-SWF	1	2	<u>OK</u>	*
2	1246	1402	S-SWF	5	2+	BE, DA, HU, MC, NE, PR, PU, SW, CW≵≴↓	1219E
3	1017	1047	S-SWF	1	2	NE	1010
3	1414	1500	S-SWF	3	2-	HU, PR	1408E
3	1757	1903	S-SWF	5	2+	AN, BE, FM, HU, LA, MC, NE, PR, WS, CW*	1756
4	0028	0208	S-SWF	5	2+	AD, OK	0032
1 1	0020	0200	5 5 11 1				0032
4	0210	0240	G-SWF	1	1+	<u>OK</u>	0208
4	0800	0840	-	1	-	CW**	*
4	1820	1920	S-SWF	5	2+	BE, FM, HU, LA, MC, PR, WS	1814
5	0620	0633	S-SWF	1	3	<u>KO</u>	*
5	1003	1022	S-SWF	3	2	NE, PU	1004E
5	1220	1232	S-SWF	5	2	BE, NE, PR, PU	1230E
5	1615	1630	S-SWF	5	1	BE, MC, PR, WS	12305
7	0440	0520	S-SWF	5	2-	CA, KE, OK	0434
7	1042	1107	S-SWF	5	1	PR, PU	*
8	0118	0200	S-SWF	5	1+	AD, CA, OK	0120E
	_						
8	0755	0820	S-SWF	5	1+	CA, KO, OK, NE	*
10	0518	0545	S-SWF	5	1+	CA, KO, <u>OK</u>	*
10	2355	0037	Slow S-SWF	5	2	AD, CA, $\overline{OK}$	*
11	0407	0430	S-SWF	1	1+	<u>OK</u>	*
12	0800	0810	-	1	-	<u>CW</u> ++	0812
17	0400	0422	Slow S-SWF	4	1+	OK, TO	*
18	0637	0657	S-SWF	4	1+	KO, OK	*
19	0345	0405	S-SWF	1	1		*
24	0343	0450	Slow S-SWF	1	2	<u>ok</u> <u>ok</u>	*
						The state of the s	

CA = Canberra, Australia DA = Darmstadt, G.F.R. JU = Juhlesruh, G.D.R.

PU = Prague, Czechoslovakia

TO = Hiraiso Radio Wave Observatory, Japan

Sudden Cosmic Noise Absorption
Sudden Enhancements Of Atmospherics
Solar Noise Bursts At 18 Mc.
SEPTEMBER 1959

				1		EMBE.			
SEPT. 1959		CLASS SEA	Burst	WIDESPREAD INDEX	(UN BEGIN	TIME IVERSAL TI MAX.	ME) END	PERCENT ABSORPTION SCNA	OBSERVATION STATIONS
1 1 1 1	3	2	1+	1 3 5 5 5	0206 1420 1655 1656 1730	0211 1425 1708 1701	0225 1428 1735 1800U 2355	15 80	HA MC, RE BO, DU, HA, NE, PA, SP BO, HA, MC, RE, SP BO, HA, MC, SP, (Noise storm, peaks at 1810-1830 (RE) 1950- 2007 (RE) 2228-2307)
1 1 2 2 2 2	2	2+ 2 1	1	3 3 5 5 5	1928 2007 1311 1605 1606 1607	1940 2022 1615 1610	2007U 2100 1329 1607 1640 1635	55	A1, A3, <u>A5</u> A1, A3, <u>A5</u> JU, KU <u>B0</u> , MC, RE, SP A3, A5, <u>B0</u> , DU, NE, PA, SP <u>B0</u> , MC, RE, SP
2 2 2 3 3	1	1	1 2 2	4 1 5 1 4	1730 1735 1800 1425 1739	1747 1752	1810 1810 0026 1605 1900	15	BO, MC BO BO, HA, MC, (Noise Storm) MC, (Group of bursts) BO, MC, SP (Group of bursts)
3 4 4 4 5		1+	1 1 2 2	1 3 5 1 5	2115 1605 1700 2320 1554	1632	0030 1710 2345 2326 1600		$\begin{array}{c} \underline{\text{HA}}  \text{(Noise Storm)} \\ \underline{\text{A2}}, \ \underline{\text{A5}} \\ \underline{\text{BO}}, \ \overline{\text{HA}}, \ \text{(Noise Storm)} \\ \underline{\underline{\text{HA}}} \\ \underline{\text{BO}}, \ \underline{\text{MC}}, \ \text{RE}, \ \text{SP} \end{array}$
6 6 7 9		2	1 2 1	1 1 1 1	2045 2158 1615 0703		2048 2205 1623 0741		HA HA MC NE
9 9 9 9	1	2	1 1 1	5 5 5 5 4	1557 1558 1620 1655 1711	1605 1611	1625 - 1625 1658 1713	20	BO, MC, RE, SP BO, MC, NE, SP BO, MC, RE, SP BO, MC, SP BO, MC
9 9 9 11 12			1 1 1 1	4 5 1 5 5	1824 1833 2235 2000 2038		1827 1839 2240 2001 2050		BO, MC BO, MC, SP MC BO, HA, MC BO, HA, MC, SP
13 14 14 14 14	1	1	1	5 1 5 5 5	1725 0745 1833 2156 2107	1740 2200	1800 0821 1835 2210 2109	10 25	BO, HA, MC NE BO, MC, SP BO, HA, SP BO, HA
16 16 19 19	1	1	1 1 1	4 5 5 5 1	1846 1846 1939 2027 2203	1850	1902 - 1948 2030 2208	15	BO, MC BO, HA, MC BO, HA, SP, (Group of bursts) BO, HA, MC, SP HA
19 20 20 21 22 23	1		1 1 1 1 1	1 5 1 1 4	2211 2147 2255 0121 1652 1909	2300 1915	2218 2148 2312 0130 1658 1917	10	HA BO, HA, MC HA HA BO, MC RE
25 25 25 26 29			1 2 1 1	1 5 4 5 5	1444 1609 1625 1830 2012		1451 1617 1635 2330 2015		MC BO, MC, RE BO, MC BO, HA, MC, SP, (Noise Storm) BO, HA
29 29			1 1	5 1	2031 2103		2035 2105		<u>во,</u> на <u>на</u>

#### SOLAR RADIO EMISSION

#### OUTSTANDING OCCURRENCES

Ottawa

JANUARY 1960

2800 Mc

Jan.	Type	Start UT	Duration	Maxim		Remarks
1960			Hrs:Mins	Time UT	Peak Flux	
3	2 Simple 2	1725	2	1725.7	12	
8	3 Simple 3	1805	2 30	1900	15	
10	1 Simple 1	1617.5	2	1618	7	
11	2 Simple 2	2056	>35	2108	220	In sunset oscillations
12	6 Complex f	1647.3	9	1649	80	
13	6 Complex f	1446	7	1449.5	18	
	4 Post Increase		15		6	į .
13	2 Simple 2	1847	4	1847.8	30	1
15	6 Complex f	b1340*	>1 40	1357	700	*In interference
	4 Post Increase A		2 30		25	1
	2 Simple 2 f	1730.5	8	1732	300	
16	3 Simple 3	ь1543	>2 17	1620	10	
17	3 Simple 3 A	ь1600	>3 30	indet.	12	
	3 Simple 3	1616	-40	1619	8	1
19	3 Simple 3 A	1925	>2 10	2007	20	
	6 Complex	1936.5	25	1945	65	
22	2 Simple 2	1649	1.5	1649.3	10	
23	1 Simple 1	1519.5	2	1520,2	6	
24	1 Simple 1	1633	1.5	1633.7	6 7	
24	1 Simple 1	1715.3	1	1715.5	7	
25	2 Simple 2	1712.5	3.5	1714.2	90	
	4 Post Increase		1 15		6	
30	3 Simple 3 f	2015	12	2021	10	

CORRECT . STANDARDS . BOW OF

#### SOLAR RADIO EMISSION

#### OUTSTANDING OCCURRENCES

JANUARY 1960

BOULDER

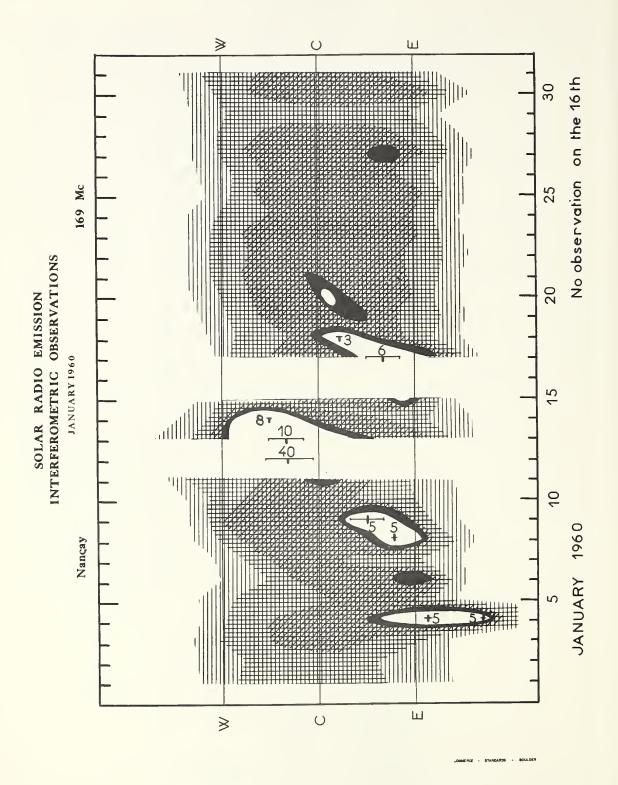
167 MC

				,	,		,	·			16/ MC
Jan. 1960	Туре	Start UT	Time of Maximum UT	Duration Minutes	Intensity	Jan. 1960	Туре	Start UT	Time of Maximum UT	Duration Minutes	Intensity
3 3 3 3	3 3 3 3	1607.1 1923.4 2133.0 2225.4 2258.3	1607.2 1923.4 2133.0 2225.6 2258.8	1.2 0.1 0.2 0.3 0.7	1 1 1 2 1**	19 19 19 20 20	7 9A 9B 2 3	1849 1941.0 1953 1432.0 1802.0	1946.2 1956.5 1433.1 1803.0	300 D 12 8 2.2 1.2	1 2 1 2* 2
3 3 4 4 4	3 3 6 3 7	2320.5 2325.8 1423 E 1806.9 1906	2320.5 2325.8 1806.9	0.1 0.2 165 D 0.1 97	1** 1** 2 1 2	20 21 21 21 21 21	3 3 3 3	1813.0 1458.0 1500.2 1602.0 1739.9	1813.0 1458.0 1500.2 1602.0 1739.9	0.1 0.1 0.1 0.2 0.1	1 1* 1* 1
5 5 6 6 7	3 3 2 2 3	1916.5 1919.6 1500.9 2319 2224.0	1916.5 1919.6 1502.8 2320 2224.0	0.1 0.2 4.1 12 0.1	2 2 1* 1**	21 22 22 22 22 24	3 3 3 3 3	2337.7 1423.0 1550.5 1805.4 2003.5	2337.7 1423.0 1550.5 1805.4 2003.5	0.2 0.1 0.2 0.2 0.3	1** 1* 2 2 2
7 7 7 8 9	3 3 6 6	2302.3 2305.4 2307.6 1423 E 1422 E	2302.3 2305.4 2308.3	0.1 0.2 1.2 554 D 458 D	1** 2** 2** 1	26 26 26 26 27	3 3 3 3 2	1422.0 1425.0 1507.2 1643.9 1417.8	1422.2 1425.0 1507.2 1643.9 1417.8	0.1 0.2 0.4 0.6 3.2	1* 1* 1 2 3*
10 10 10 10 10	3 3 3 3 7	1836.0 1842.3 2021.5 2143.7 2229	1837.1 1842.3 2021.5 2143.7	2.0 0.3 0.1 0.2 71 D	2 2 1 2 2	27 27 27 27 27 27	3 3 3 3 3	1428.0 1504.7 1531.3 1543.0 1805.5	1428.0 1504.7 1531.3 1543.0 1806.0	0.2 0.5 0.4 0.8 0.7	2* 2 1 1 2
11 11 12 12 13	6 9 6 8 6	1421 E 2056.0 1422 E 1648.9 1421 E	1935 U 1651.2 1957 U	562 D 167 D 561 D 12 562 D	2 3** 2 3 1	27 27 27 27 27 28	3 3 3 3	1809.0 1838.5 1852.0 2348.8 1432.5	1809.0 1839.0 1852.5 2348.9 1432.5	0.2 1.2 1.0 0.8 0.5	2 1 1 2** 2*
14 15 16 16 17	6 6 8 6	1421 E 1418 E 1422 E 2247 1421 E	1756 U	564 D 567 D 565 D 6 566 D	2 1 2 3 1	28 29 30 30 30	3 3 3 2	2334.5 1823.2 1546.0 1640.0 1714	2334.5 1823.2 1546.0 1640.3 1714.0	0.3 0.2 1.0 0.4 7	2** 2 2 2 2
17 18 18 18 18	2 2 2 3 3	1910.0 1705.0 1739 1757.0 1912.8	1911.0 1705.6 1740 1757.0 1912.8	1.6 1.0 7 0.2 0.3	2 2 2 2 2	30 30 30 30 30	3 3 3 3	1723.0 1836.0 2116.6 2201.0 2348.6	1723.0 1836.0 2116.6 2201.0 2348.6	0.5 0.9 0.4 0.2 0.3	3 1 2 3 1**
18 19 19 19 19	3 3 3 3	2330.2 1431.6 1433.8 1455.8 1640.0	2330.2 1431.7 1433.8 1455.8 1640.0	0.2 0.8 0.1 0.3 0.1	2** 2* 2* 2 1	31 31 31 31 31	3 3 3 3	1423.0 1451.2 2002.0 2350.9 2354.1	1423.0 1451.2 2002.0 2351.2 2355.2	0.2 0.1 0.8 1.6 2.8	1* 2* 2 2** 2**

#### TIMES OF OBSERVATION

Jan. 1960	U.T.	Jan. 1960	
1 2 3 4 5 6 7 8 9 10 11 12	1422-2329 1424-2333 1424-2334 1423-2335 1422-2337 1423-2337 1423-2337 1423-2337 1421-2340 1421-2343 1421-2343 1421-2343 1421-2343 1421-2345 1421-2345	16 17 18 19 20 21 22 23 24 25 26 27 28 29 30 31	 I after 1830 I 1915-2030 I throughout day I throughout day I 1830-0004 I 1725-1745

<sup>\*</sup> On sunrise pattern \*\* On sunset pattern

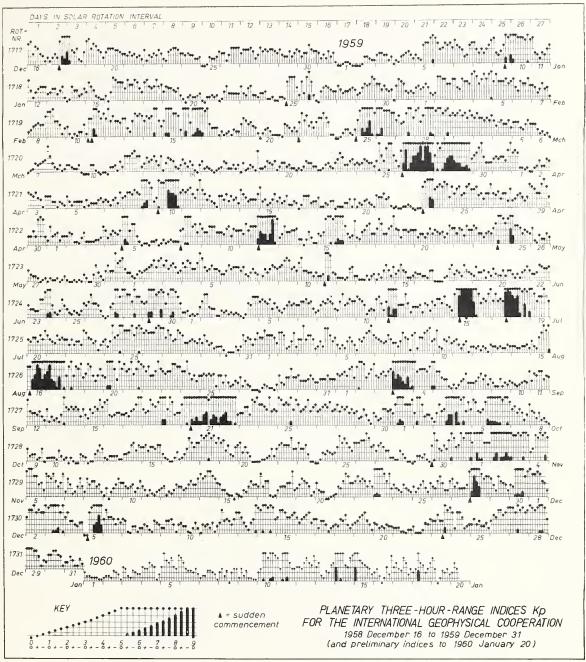




#### GEOMAGNETIC ACTIVITY INDICES

DECEMBER 1959

Dec. 1959	С	Values Three hour Gr 1 2 3 4		Sum	Ар	Final Selected Days
1 2 3 4 5	1.1 1.2 1.5 0.7 1.7	50 4+ 4+ 5- 3+ 4+ 4+ 50 50 5- 50 5+ 3+ 40 3- 3- 00 00 40 5-	4+ 3+ 30 2- 40 4+ 40 40 5+ 6- 40 5+ 20 2- 2+ 2+ 6+ 7- 8- 6+	31- 33+ 40+ 210 36-	28 30 50 12 68	Five Quiet 7 10 11
7 8 9 10	0.9 0.3 0.3 0.3 0.2	30 4+ 3+ 3+ 30 3- 1- 0+ 1+ 20 2- 10 1+ 1+ 3- 3- 2+ 2- 2- 1-	30 20 30 2+ 2- 2- 1- 1+ 1- 3- 2+ 20 30 20 1- 20 0+ 1+ 2+ 1+	24+ 120 14- 16- 12-	16 6 7 8 6	21 22
11 12 13 14 15	0.2 1.0 1.1 1.4 1.1	0+ 20 2- 20 2+ 4+ 3- 30 40 3+ 2+ 2- 5- 4+ 5+ 50 3+ 4+ 4+ 40	1+ 2+ 1- 2- 3- 20 30 40 10 2+ 30 5+ 5- 5+ 5- 30 3+ 3+ 40 4+	12o 24o 23o 37o 31o	6 16 18 40 26	Five Disturbed 3 5 14
16 17 18 19 20	1.0 0.3 0.5 0.8 0.3	3+ 2+ 30 3+ 30 1+ 3- 3- 20 2- 30 3- 2+ 3- 3+ 4- 2+ 20 1+ 2-	4- 40 2+ 30 20 1- 1+ 2- 4- 20 1+ 20 40 3- 20 10 1+ 20 2- 2-	250 15+ 18+ 22- 140	17 8 10 14 6	27 28
21 22 23 24 25	0.1 0.4 1.3 0.9 0.4	1+ 1+ 1+ 1- 2- 0+ 0+ 0+ 3- 3- 4- 3- 40 1+ 40 40 20 30 20 3-	1- 0+ 10 10 10 2+ 20 30 3- 60 4+ 5- 3+ 1+ 2- 20 2- 2- 2+ 1+	8- 11o 29+ 22- 17-	4 6 28 15 8	Ten Quiet 7 8 9 10
26 27 28 29 30 31	1.2 1.4 1.4 0.9 0.8 0.4	3- 4- 3+ 4+ 5- 5- 40 6- 50 5- 4+ 5- 40 40 40 4- 2- 30 30 4- 2- 2- 3- 20	4+ 4- 40 5- 4+ 5- 40 4+ 5+ 5+ 5- 30 4- 3- 10 20 4- 30 3- 10 10 10 2+ 1+	31- 36+ 370 250 22- 14-	26 38 40 18 14 7	11 17 20 21 22 31
Mean:	0.81			Mean:	19	

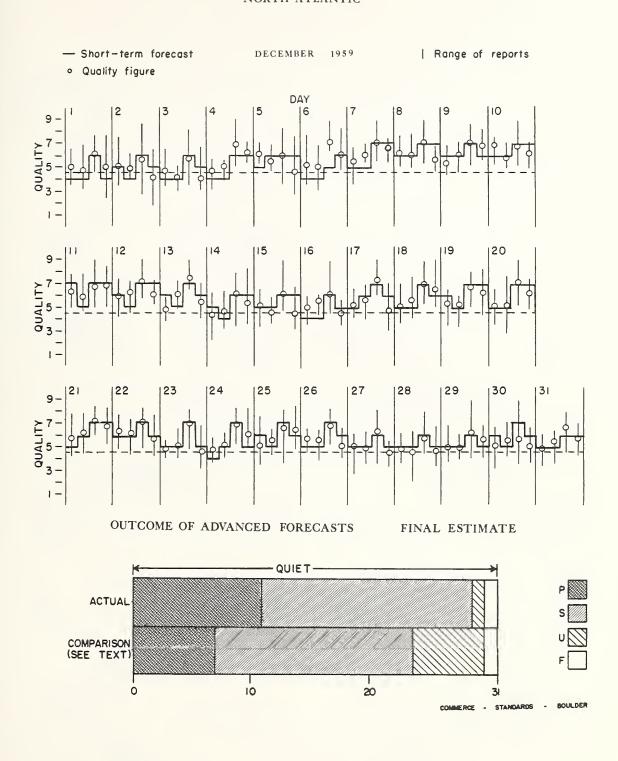


## CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH ATLANTIC

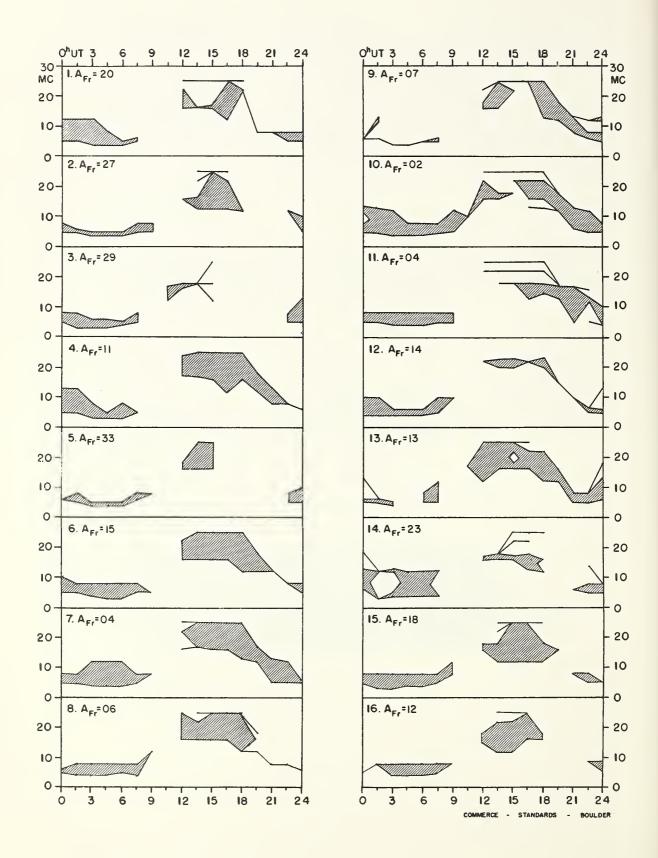
DECEMBER 1959

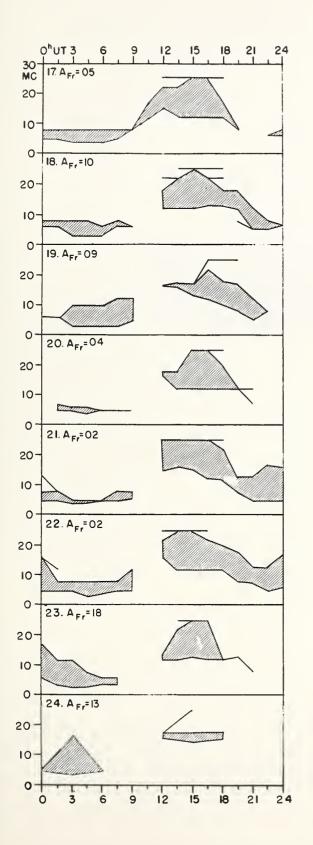
Dec. 1959	North Atlantic 6-hourly quality figures 00 06 12 18 to to to to 06 12 18 24	Short-term forecasts issued about one hour in advance of:	Whole day index	Advance forecasts (J-reports) for whole day; issued in advance by: 1-7 1-7 1-7 1-7 days days days Final Js SDW J	Geomag- netic K <sub>Fr</sub> Half Day (1) (2)
1 2 3 4 5	50 5- 6+ 50 5+ 50 6- 40 5- 4+ 6- 4+ 5- 50 70 6+ 6+ 6- 60 5-	4 4 6 4 5 4 6 5 4 4 6 5 4 4 6 6 5 6 6 6	5+ 5- 5- 6- 6-	4 4 5 5 4 5 5 5 5 5 5	(4) 3 (4) (4) (5) (4) 3 1 2 (5)
6 7 8 9 10	5+ 50 70 6+ 6- 60 70 7- 6+ 6+ 7+ 6- 5+ 6+ 70 70 7- 60 70 6+	4 4 5 6 5 5 7 7 6 6 7 7 6 6 7 6 6 6 7 7	60 6+ 6+ 6+ 6+	6 6 7 7 7 7 7 7 6 6 6	3 2 1 1 1 2 2 2 1 1
11 12 13 14 15	6+ 6- 7- 7- 60 6+ 70 60 5- 60 7+ 5+ 4+ 5- 60 5+ 5+ 4+ 6+ 4+	7 5 7 7 6 5 7 7 6 5 7 6 5 4 6 6 5 5 6 6	6+ 6+ 60 50 50	6 6 7 7 7 7 7 7 6 6 6	1 1 3 3 3 3 3 (4) (4) 3
16 17 18 19 20	50 5+ 60 4+ 50 6- 7+ 5- 50 6- 70 7- 5+ 5+ 7- 6+ 50 5+ 7+ 6+	4 4 6 5 5 6 7 6 5 5 7 6 6 5 7 7 5 5 7 7	50 6- 60 6- 60	6 6 6 6 6 6 5 5	3 2 2 1 2 2 3 1 2 1
21 22 23 24 25	6- 60 70 7- 6+ 60 70 6- 5- 50 70 4+ 5- 50 7- 60 50 6- 7- 6+	5 6 7 7 6 6 7 6 5 5 7 5 4 5 7 5 6 5 7 6	6+ 6+ 5+ 6-	5 5 6 6 6 6 4 4 5 5	1 1 0 1 2 (4) 3 2 2 1
26 27 28 29 30 31	6- 6- 7- 50 50 50 6+ 4+ 5- 4+ 6- 5- 50 50 6+ 6- 50 6- 6- 50 5- 5+ 7- 6-	5 5 7 6 5 5 6 5 5 5 6 5 5 5 6 5 6 5 7 6 5 5 6 6	6- 50 5- 6- 5+ 6-	6 6 6 6 6 6 6 6 6 6 6	(4) 3 (4) (4) (4) (4) (4) 2 3 2 2 1
Score	: Quiet Periods	P 13 14 27 9 S 17 14 3 16 U 0 0 1 0 F 0 0 0 0		11 11 18 18 1 1 1 1	
	isturbed Periods	P 0 1 0 0 S 1 2 0 5 U 0 0 0 0 F 0 0 0 1		0 0 0 0 0 0 0 0	

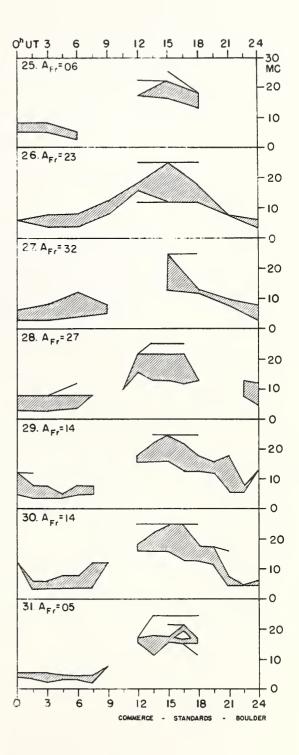
( ) represent disturbed values.



DECEMBER 1959







#### CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS

#### NORTH PACIFIC

DECEMBER 1959

Dec. 1959	North F 12-hou quality	rly		erm fore- ssued at	Whole day index	(Jp repo	forecasts orts) for ny; issued vance by:	Geomag- netic K <sub>Si</sub>
	0700 to 1900	1900 to 0700	0600	1800		l-7 l-7 days day Final Jps	s days days	Half Day (1) (2)
1 2 3 4 5	6 6 4 6 3	6 6 4 5 2	4 4 5 5 5	5 5 5 5	6 5 (4) 5 (2)	4 5 5 5 6	4 5 5 5 6	(5) (4) (4) (4) (5) (6) 3 2 2 (8)
6 7 8 9 10	6 5 5 6 5	6 5 6 5	5 6 5 5	6 5 5 6 5	6 5 5 6 5	6 6 6 6 5	6 6 6 5	3 2 1 2 1 1 2 1 1 1
11 12 13 14 15	5 7 6 5 5	4 5 6 5	5 5 5 5 4	5 5 5 5 5	5 6 6 (4) 5	5 6 6 6	5 6 6 6	1 1 2 2 2 2 2 2 (5) (4) (4) (4)
16 17 18 19 20	5 6 7 6 6	5 6 7 6 5	5 5 5 5 5	5 6 6 6	5 6 7 6 5	5 5 5 6 5	5 5 6 5	3 (4) 2 2 2 2 (4) 3 1 2
21 22 23 24 25	6 5 5 6 6	6 6 6 5 7	5 5 5 5 5	5 5 6 6	6 6 6 6	5 6 6 5 6	5 6 6 5	0 0 0 2 2 (4) (4) 2 2 2
26 27 28 29 30 31	5 5 4 5 6 6	6 5 5 6 6	6 5 4 5 5	5 5 4 5 6	6 5 (4) 5 6 6	6 5 5 5 5 6	6 5 5 5 6	(4) (4) (4) (4) (4) (5) (4) 3 2 3 1 2
Score:	Quiet	Periods	P 10 S 14 U 2 F 2	14 14 0 0		18 7 1		
1	Disturbed	Periods	P 1 S 1 U 1 F 0	0 2 0 1		0 2 0 2		

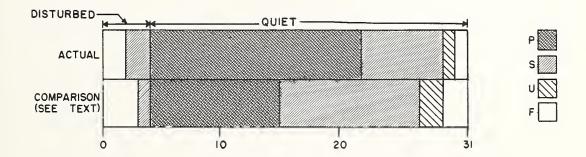
( ) represent disturbed values.

COMMERCE - STANDARDS - BOULDER

Errata: In CRPL-F 185 Part B the score for December 1959 under "Whole day index" should have been under "1-7 days Final".

NORTH PACIFIC DECEMBER 1959

OUTCOME OF ADVANCED FORECASTS FINAL ESTIMATE



### INTERNATIONAL WORLD DAY SERVICE JANUARY 1960

Issued Day/Time UT Jan 1960	Advance Geophysical Alert	No.	World-Wide Geophysical Alert	Special World Interval
11/1600		44	Magnetic Storm 10/0719Z	
12/0000	Burbank Solar Flare 11/2140Z			
14/1005	Fort Belvoir Magnetic Storm 13/1900Z			(
14/1600		45	Magnetic Storm 13/1900Z	
18/1115	Fort Belvoir Magnetic Storm 17/12XXZ			
18/1600		46 (	Magnetic Storm 17/1200Z	
21/1600		47	Magnetic Storm 21/00XXZ	



